# $T \bar{T}$ deformations of holographic warped CFTs 

Rahul Poddar©*<br>Science Institute, University of Iceland, Dunhaga 3, 107 Reykjavík, Iceland

(Received 21 September 2023; accepted 7 November 2023; published 27 November 2023)


#### Abstract

We explore $T \bar{T}$ deformations of warped conformal field theories (WCFTs) in two dimensions as examples of $T \bar{T}$ deformed nonrelativistic quantum field theories. WCFTs are quantum field theories with a Virasoro $\times \mathrm{U}(1)$ Kac-Moody symmetry. We compute the deformed symmetry algebra of a $T \bar{T}$ deformed holographic WCFT, using the asymptotic symmetries of $\mathrm{AdS}_{3}$ with $T \bar{T}$ deformed Compére, Song, and Strominger boundary conditions. The $\mathrm{U}(1)$ Kac-Moody symmetry survives provided one allows the boundary metric to transform under the asymptotic symmetry. The Virasoro sector remains but is now deformed and no longer chiral.


DOI: 10.1103/PhysRevD.108.105016

## I. INTRODUCTION

A warped conformal field theory (WCFT) is a quantum field theory with an $\operatorname{SL}(2, \mathbb{R}) \times \mathrm{U}(1)$ global symmetry in two dimensions, which breaks Lorentz invariance. Such QFTs have translation invariance, but scaling invariance is restricted to only one coordinate. Finite warped symmetry transformations take the form [1,2]

$$
\begin{equation*}
z \rightarrow f(z), \quad \bar{z} \rightarrow \bar{z}+g(z) \tag{1.1}
\end{equation*}
$$

However, despite not being Lorentz invariant, this class of two-dimensional quantum field theories still possesses an infinite-dimensional symmetry algebra, namely, a Virasoro $\times \mathrm{U}(1)$ Kac-Moody current algebra. WCFTs are interesting, as they appear in a number of holographic systems with an $\operatorname{SL}(2, \mathbb{R}) \times \mathrm{U}(1)$ symmetry, such as warped $\mathrm{AdS}_{3}$ [3], the near-horizon geometry of extremal rotating black holes $[4,5]$, and $\mathrm{AdS}_{3}$ with DirichletNeumann boundary conditions [6]. Holographic WCFTs have passed a number of consistency checks, such as a Cardy formula [2], holographic entanglement entropy [7-10], and one-loop determinants [11].

Since WCFTs are nonrelativistic, they do not couple to standard (pseudo-)Riemannian manifolds. One approach is to couple WCFTs to "warped geometries" [3], a variant of Newton-Cartan geometries. These geometries can be found at the boundary of warped $\mathrm{AdS}_{3}$ spacetimes. Unfortunately, these geometries have certain pathologies,

[^0]Published by the American Physical Society under the terms of the Creative Commons Attribution 4.0 International license. Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI. Funded by SCOAP ${ }^{3}$.
such as a degenerate metric, which make some calculations untenable. Another way to couple WCFT to a background manifold is to allow the manifold to transform with the warped symmetry transformations. Holographically, this requires relaxing Dirichlet boundary conditions of the bulk metric to boundary conditions which allow for asymptotic symmetry transformations to transform the boundary metric under the warped symmetry transformation of the boundary WCFT. The Dirichlet-Neumann boundary conditions of Compére, Song, and Strominger (CSS) [6] do exactly this. This approach bypasses the need for a warped geometry with degenerate metrics, and we can work with conventional techniques.

Two-dimensional translationally invariant quantum field theories admit a class of solvable irrelevant deformations built from conserved currents, most notable of which is the $T \bar{T}$ deformation $[12,13]$. The $T \bar{T}$ operator is defined by the determinant of the energy-momentum tensor of the quantum field theory, and the deformed action obeys the following flow equation:

$$
\begin{align*}
\partial_{\lambda} S_{\mathrm{QFT}}(\lambda) & =-\frac{1}{2} \int d^{2} x \sqrt{\gamma} \mathcal{O}_{T \bar{T}}^{(\lambda)} \\
\mathcal{O}_{T \bar{T}} & =\operatorname{det} T=\frac{1}{2} \epsilon^{\mu \rho} \epsilon^{\nu \sigma} T_{\mu \nu} T_{\rho \sigma} \tag{1.2}
\end{align*}
$$

where the deformation parameter $\lambda$ is the coupling to the $T \bar{T}$ operator $\mathcal{O}_{T \bar{T}}$. This operator is defined using point splitting, which in the coincident limit produces a welldefined local operator up to total derivatives. The expectation value of $\mathcal{O}_{T \bar{T}}$ turns out to be a constant, and from this one can derive the flow of energy eigenstates of the quantum field theory defined on a cylinder of radius $R$ :

$$
\begin{equation*}
\frac{\partial E_{n}}{\partial \lambda}=E_{n} \frac{\partial E_{n}}{\partial R}+\frac{P_{n}^{2}}{R} \tag{1.3}
\end{equation*}
$$

Even though the energy eigenvalues are changed, the Hilbert space remains undeformed, since there is a one-one correspondence between the states of the original and deformed theory. Similarly, other observables can be calculated in the deformed theory, for example, the deformed Lagrangian, partition function, two-two scattering matrices, correlation functions, etc. [14-21].

There are various ways to interpret how the $T \bar{T}$ deformation acts on holographic CFTs. One proposal by [22] is to impose Dirichlet boundary conditions for the bulk metric at a finite radius. Another proposal given by [23] is to treat the $T \bar{T}$ deformation as a double-trace deformation, which will deform the asymptotic behavior of the bulk fields [24,25]. This approach agrees with the cutoff AdS proposal when both are valid but has the advantage of working when there are bulk matter fields and also for either sign of the deformation parameter, which the former does not. More recently, there has also been the "glue-on AdS holography" proposal [26] which also agrees with [23] for the positive sign of the deformation parameter in the absence of matter fields.

Using the mixed boundary conditions and Dirichlet boundary conditions at finite radius, the asymptotic symmetry algebra of the bulk dual to $T \bar{T}$ deformed holographic CFT was calculated in [23,27,28]. Despite losing conformal invariance, the asymptotic symmetry algebra turns out to still have a Virasoro $\times$ Virasoro structure. However, either the central charge becomes state dependent, or one loses the holomorphic factorization of the symmetry algebra, which can also be expressed as a nonlinear deformation of the standard Virasoro algebra.

In this work, we explore $T \bar{T}$ deformations of WCFTs from a holographic perspective. To establish what a $T \bar{T}$ deformation of a WCFT is, one must first define what the energy-momentum tensor of a WCFT is. Canonically, for WCFTs defined on a warped geometry, energy-momentum tensors are not symmetric, and a determinant is harder to define, since the metric is degenerate and noninvertible. The energy-momentum tensor turns out to be a tensor with components being a chiral stress tensor and a $\mathrm{U}(1)$ current. For WCFTs dual to warped $\mathrm{AdS}_{3}$, it is also not possible to use the Fefferman-Graham expansion to compute the energy-momentum tensor for the same reason; the boundary metric is not invertible. However, if we study WCFTs dual to $\mathrm{AdS}_{3}$ with CSS boundary conditions, for the price of a boundary metric which is not invariant under warped transformations, we have an invertible metric, a symmetric energy-momentum tensor, and a conventional definition for a determinant. Given these considerations, it is possible to propose a definition for a $T \bar{T}$ deformed WCFT which is dual to $\mathrm{AdS}_{3}$ with $T \bar{T}$ deformed CSS boundary conditions.

This paper is organized as follows. We first briefly review the mixed boundary conditions of [23] in Sec. II A. Then, in Sec. II B, we review the CSS boundary conditions and derive the Virasoro $\times \mathrm{U}(1)$ Kac-Moody algebra. In Sec. III, we derive the $T \bar{T}$ deformed CSS boundary
conditions to compute the deformed symmetry algebra of a $T \bar{T}$ deformed WCFT. We will see that if one imposes deformed boundary conditions equivalent to Dirichlet boundary conditions at the radial cutoff surface, we recover a deformed Virasoro algebra, but we lose the $\mathrm{U}(1)$ KacMoody algebra. However, if we allow the DirichletNeumann boundary conditions to remain at the cutoff surface, which is what the mixed boundary conditions suggests is the correct approach, we recover an undeformed Kac-Moody symmetry. We then conclude and discuss future directions in Sec. IV.

## II. REVIEW

## A. Mixed boundary conditions from $\boldsymbol{T} \bar{T}$

We begin by briefly reviewing the mixed boundary conditions derived in [23] from the variational principle. The variation of the boundary QFT action with respect to the boundary metric sources the energy-momentum tensor of the QFT and of the bulk dual. The flow of the variation of the QFT action is equal to the variation of the deformation which generates the flow. So we have

$$
\begin{align*}
\partial_{\lambda} \delta S & =\delta \partial_{\lambda} S \\
\partial_{\lambda}\left(\frac{1}{2} \int_{\partial \mathcal{M}} d^{2} x \sqrt{\gamma} T_{i j}^{(\lambda)} \delta \gamma^{(\lambda) i j}\right) & =\delta \int_{\partial \mathcal{M}} d^{2} x \sqrt{\gamma} \mathcal{O}_{T \bar{T}}^{(\lambda)} . \tag{2.1}
\end{align*}
$$

From this, we can compute the flow equations for the boundary metric and the energy-momentum tensor with respect to the deformation parameter $\lambda$. Expressing the equations in terms of the trace-reversed energy-momentum tensor $\hat{T}_{i j}=T_{i j}-\gamma_{i j} T_{i}^{i}$, we have

$$
\begin{align*}
\partial_{\lambda} \gamma_{i j} & =-2 \hat{T}_{i j}, \\
\partial_{\lambda} \hat{T}_{i j} & =-\hat{T}_{i l} \hat{T}_{j}^{l}, \\
\partial_{\lambda}\left(\hat{T}_{i l} \hat{T}_{j}^{l}\right) & =0 . \tag{2.2}
\end{align*}
$$

Solving these equations, we can express the deformed metric and energy-momentum tensor in terms of the undeformed metric and energy-momentum tensor:

$$
\begin{align*}
& \gamma_{i j}(\lambda)=\gamma_{i j}-2 \lambda \hat{T}_{i j}+\lambda^{2} \hat{T}_{i k} \hat{T}_{j l} \gamma^{k l} \\
& \hat{T}_{i j}(\lambda)=\hat{T}_{i j}-\lambda \hat{T}_{i k} \hat{T}_{j l} \gamma^{k l} \tag{2.3}
\end{align*}
$$

where everything on the right-hand side is undeformed quantities. The new deformed quantities are now the new boundary conditions for the bulk fields. To see this, let us consider pure Einstein gravity.

For pure Einstein gravity in three dimensions, the Fefferman-Graham expansion of the metric truncates at second order in $1 / r^{2}$ [29]:

$$
\begin{align*}
d s^{2} & =l^{2} \frac{d r^{2}}{r^{2}}+g_{a b} d z^{a} d z^{b} \\
& =l^{2} \frac{d r^{2}}{r^{2}}+l^{2} r^{2}\left(g_{a b}^{(0)}+\frac{g_{a b}^{(2)}}{r^{2}}+\frac{g_{a b}^{(4)}}{r^{4}}\right) d z^{a} d z^{b} \tag{2.4}
\end{align*}
$$

using which we can now express the boundary energymomentum tensor in terms of the Fefferman-Graham expansion:

$$
\begin{equation*}
\hat{T}_{a b}=\frac{k}{2 \pi} g_{a b}^{(2)} \tag{2.5}
\end{equation*}
$$

where $k=\frac{l}{4 G_{N}}$. For pure gravity, we also have

$$
\begin{equation*}
g_{a b}^{(4)}=\frac{1}{4} g_{a c}^{(2)} g_{d b}^{(2)} g_{(0)}^{c d} \tag{2.6}
\end{equation*}
$$

Therefore, we can express the deformed boundary metric and energy-momentum tensor in terms of the FeffermanGraham expansion:
$\gamma_{a b}(\lambda)=l^{2}\left(g_{a b}^{(0)}-\left(2 \lambda \frac{k}{2 \pi}\right) g_{a b}^{(2)}+\left(2 \lambda \frac{k}{2 \pi}\right)^{2} g_{a b}^{(4)}\right)$,
$\hat{T}_{a b}(\lambda)=\frac{k}{2 \pi}\left(g_{a b}^{(2)}-\left(2 \lambda \frac{k}{2 \pi}\right) g_{a b}^{(4)}\right)$.
Equating this to the Fefferman-Graham expansion (2.4), it is easy to see that the deformed boundary metric can be thought of as being placed at a finite radius $r_{c}=\sqrt{-\frac{\pi}{k \lambda}}$. Indeed, it turns out that the Brown-York energy-momentum tensor (with the appropriate counterterm) evaluated at this surface reproduces the deformed energy-momentum tensor derived here. This makes it clear that, in pure gravity, the mixed boundary conditions and imposing Dirichlet boundary conditions at $r_{c}=\sqrt{-\frac{\pi}{k \lambda}}$ are equivalent. ${ }^{1}$

For a derivation of the mixed boundary conditions from the Chern-Simons formulation of 3D gravity, see Ref. [30].

## B. CSS boundary conditions

Examples of constructing a holographic bulk dual to a WCFT are either warped $\mathrm{AdS}_{3}$ or $\mathrm{AdS}_{3}$ with CSS boundary conditions. We shall use the CSS boundary conditions, since they are amenable to the mixed boundary conditions from the $T \bar{T}$ deformation.

Expressing the metric in Fefferman-Graham gauge (2.4), we have the following Dirichlet-Neumann boundary conditions for the metric [6]:

$$
g^{(0)}=\left(\begin{array}{cc}
P^{\prime}(z) & -\frac{1}{2}  \tag{2.8}\\
-\frac{1}{2} & 0
\end{array}\right), \quad g_{\bar{z} \bar{z}}^{(2)}=\frac{\Delta}{k},
$$

${ }^{1}$ We will be absorbing the factor of $\pi$ into the normalization of
$\lambda$ and $\mathcal{O}_{T \bar{T}}$ from now on to avoid clutter in the equations.
where $k=\frac{l}{4 G_{N}}$ and $\Delta$ is a constant. These falloff conditions are chiral, with $P(z)$ being an undetermined holomorphic function. This is to accommodate (1.1), which shifts $P^{\prime}(z)$, and, hence, we must leave it undetermined. Note that this is unlike the warped geometry in [3], where the warped geometry metric is invariant under (1.1).

One can compute the full bulk metric with the CSS boundary conditions by taking the Fefferman-Graham expansion (2.4) to be

$$
\begin{align*}
\frac{d s^{2}}{l^{2}}= & \frac{d r^{2}}{r^{2}}+\frac{\Delta}{k} d \bar{z}^{2}-\left(r^{2}+\frac{2 \Delta P^{\prime}(z)}{k}+\frac{\Delta L(z)}{k^{2} r^{2}}\right) d z d \bar{z} \\
& +\left(r^{2} P^{\prime}(z)+\frac{\left(L(z)+\Delta P^{\prime}(z)\right)^{2}}{k}+\frac{\Delta L(z) P^{\prime}(z)}{k^{2} r^{2}}\right) d z^{2} \tag{2.9}
\end{align*}
$$

Here, both $L(z)$ and $P(z)$ are undetermined holomorphic functions and parametrize the phase space of $\mathrm{AdS}_{3}$ with CSS boundary conditions. A special case is the Banãdos-Teitelboim-Zanelli black hole when $P^{\prime}(z)=L^{\prime}(z)=0$ [6,31].

The asymptotic symmetries of this metric are interesting, as they differ from the usual product of $\operatorname{SL}(2, \mathbb{R})$ algebras despite being locally $\mathrm{AdS}_{3}$. To compute asymptotic Killing vectors, we first require that they preserve radial gauge:

$$
\begin{equation*}
\mathcal{L}_{\xi} g_{r \mu}=0 \tag{2.10}
\end{equation*}
$$

This fixes the asymptotic Killing vector $\xi$ to have the form
$\xi=r f(z, \bar{z}) \partial_{r}+\left(V^{a}(z, \bar{z})-\int \frac{g^{a b}}{r} \partial_{b} f(z, \bar{z}) d r\right) \partial_{a}$.

Evaluating this for the CSS metric (2.9), we get
$\xi^{r}=r f(z, \bar{z})$,

$$
\begin{align*}
\xi^{z}= & V^{z}(z, \bar{z})-\frac{k\left(\partial_{z} f(z, \bar{z})+\left(k r^{2}+\Delta P^{\prime}(z)\right) \partial_{\bar{z}} f(z, \bar{z})\right)}{k^{2} r^{4}-\Delta L(z)}, \\
\xi^{\bar{z}}= & V^{\bar{z}}(z, \bar{z})-\frac{k}{k^{2} r^{4}-\Delta L(z)}\left(\left(k r^{2}+\Delta P^{\prime}(z)\right) \partial_{z} f(z, \bar{z})\right. \\
& \left.+\left(\left(2 k r^{2}+\Delta P^{\prime}(z)\right) P^{\prime}(z)+L(z)\right) \partial_{\bar{z}} f(z, \bar{z})\right) . \tag{2.12}
\end{align*}
$$

If we impose Dirichlet boundary conditions at infinity,

$$
\begin{equation*}
\lim _{r \rightarrow \infty} \mathcal{L}_{\xi} g_{\mu \nu}=0 \tag{2.13}
\end{equation*}
$$

we get conditions on the undetermined functions in $\xi$ :

$$
\begin{align*}
\partial_{\bar{z}} V^{a}(z, \bar{z}) & =0, \quad f(z, \bar{z})=-\frac{1}{2} \partial_{z} V^{z}(z, \bar{z}), \\
V^{\bar{z}}(z, \bar{z}) & =P^{\prime}(z) V^{z}(z, \bar{z}) \tag{2.14}
\end{align*}
$$

and so we can write our asymptotic killing vector [where $\left.V(z) \equiv V^{z}(z, \bar{z})\right]$

$$
\begin{align*}
\xi(V)= & -\frac{1}{2} V^{\prime}(z) \partial_{r}+\left(V(z)+\frac{k \Delta V^{\prime \prime}(z)}{2\left(k^{2} r^{4}-\Delta L(z)\right)}\right) \partial_{z} \\
& +\left(P^{\prime}(z) V(z)+\frac{k\left(k r^{2}+\Delta P^{\prime}(z)\right)}{2\left(k^{2} r^{4}-\Delta L(z)\right)} V^{\prime \prime}(z)\right) \partial_{\bar{z}} \tag{2.15}
\end{align*}
$$

Asymptotic Killing vectors generate flow in the phase space; i.e.,

$$
\begin{equation*}
\mathcal{L}_{\xi} g_{\mu, \nu}=\partial_{L(z)} g_{\mu \nu} \delta_{\xi} L(z)+\partial_{P^{\prime}(z)} g_{\mu \nu} \delta_{\xi} P^{\prime}(z) \tag{2.16}
\end{equation*}
$$

From this, we can compute $\delta L$ and $\delta P$. It turns out that $\xi$ transforms only $L(z)$ and reproduces the infinitesimal Schwarzian transformation:
$\delta_{\xi} L(z)=V(z) L^{\prime}(z)+2 V^{\prime}(z) L(z)-\frac{k}{2} V^{\prime \prime \prime}(z), \quad \delta_{\xi} P=0$.

To transform $P(z)$, we cannot allow the asymptotic killing vector to satisfy Dirichlet boundary conditions at infinity (2.13), since warped symmetry requires changing the boundary metric. The "asymptotic Killing vector" ${ }^{2}$ which generates transformations in $P(z)$ is

$$
\begin{equation*}
\eta(\sigma)=\sigma(z) \partial_{\bar{z}}, \tag{2.18}
\end{equation*}
$$

and the transformations of the parametrizing functions are

$$
\begin{equation*}
\delta_{\eta} L=0, \quad \delta_{\eta} P(z)=-\sigma(z) \tag{2.19}
\end{equation*}
$$

Note that $\eta$ also generates the warped symmetry transformation $\bar{z} \rightarrow \bar{z}+\sigma(z)$.

We can use the Fefferman-Graham expansion to compute the boundary energy-momentum tensor

$$
\begin{align*}
T_{a b} & =\frac{k}{2 \pi}\left(g_{a b}^{(2)}-g_{(0)}^{k l} g_{k l}^{(2)} g_{a b}^{(0)}\right) \\
& =\frac{1}{2 \pi}\left(\begin{array}{cc}
L(z)+\Delta P^{\prime}(z)^{2} & -\Delta P^{\prime}(z) \\
-\Delta P^{\prime}(z) & \Delta
\end{array}\right) . \tag{2.20}
\end{align*}
$$

At this point, it should be stated that this energy-momentum tensor is not the canonical energy-momentum tensor for a warped CFT. For a warped CFT defined on a manifold with warped geometry, the energy-momentum tensor is not

[^1]symmetric, since symmetry of the energy-momentum tensor is a result of Lorentz invariance. However, a warped CFT dual to $\mathrm{AdS}_{3}$ with CSS boundary conditions is not defined on a manifold with warped geometry. Rather, the manifold is not invariant under warped transformations, but for that price we gain the symmetry of the energymomentum tensor.

The conserved charges corresponding to the asymptotic Killing vectors are
$Q_{\xi(f)}=\frac{1}{2 \pi} \int_{\partial \Sigma} d \phi n^{a} T_{a b} \xi^{b}=\frac{1}{4 \pi^{2}} \int_{0}^{2 \pi} d \phi f(z) L(z)$,
$Q_{\eta(f)}=\frac{1}{2 \pi} \int_{\partial \Sigma} d \phi n^{a} T_{a b} \eta^{b}=\frac{\Delta}{4 \pi^{2}} \int_{0}^{2 \pi} d \phi f(z)\left(P^{\prime}(z)-1\right)$,
where $\partial \Sigma$ is at $r \rightarrow \infty, t=\frac{z+\bar{z}}{2}$ constant, $\phi=\frac{z-\bar{z}}{2} \in(0,2 \pi)$, and $n=\partial_{t}=\partial_{z}+\partial_{\bar{z}}$.

We can now also compute the charge algebra, using the Dirac brackets of Einstein gravity:

$$
\begin{equation*}
\left\{Q_{\zeta_{1}(f)}, Q_{\zeta_{2}(g)}\right\}=\delta_{\zeta_{1}(f)} Q_{\zeta_{2}(g)} \tag{2.22}
\end{equation*}
$$

So we have

$$
\begin{align*}
\left\{Q_{\xi(f)}, Q_{\xi(g)}\right\}= & \delta_{\xi(f)} Q_{\xi(g)}=\frac{1}{4 \pi^{2}} \int_{0}^{2 \pi} d \phi g(z) \delta_{\xi(f)} L(z) \\
= & \frac{1}{4 \pi^{2}} \int_{0}^{2 \pi} g(z) \\
& \times\left(f(z) L^{\prime}(z)+2 f^{\prime}(z) L(z)-\frac{k}{2} f^{\prime \prime \prime}(z)\right) \tag{2.23}
\end{align*}
$$

Expanding the functions in modes,

$$
\begin{align*}
& f(z)=\sum_{n} f_{n} e^{i n z}, \quad g(z)=\sum_{m} g_{m} e^{i m z}, \\
& L(z)=\sum_{p} L_{p} e^{-i p z} \tag{2.24}
\end{align*}
$$

Replacing Dirac brackets with commutators, we obtain the Virasoro algebra

$$
\begin{equation*}
\left[L_{m}, L_{n}\right]=(m-n) L_{m+n}-\frac{k}{2} n^{3} \delta_{m,-n} \tag{2.25}
\end{equation*}
$$

Note that equating $\frac{k}{2}=\frac{c}{12}$ gives the familiar $c=6 k=\frac{3 l}{2 G_{N}}$. Similarly, we obtain a U(1) Kac-Moody algebra from the commutator of the charges $Q_{\eta}$ :

$$
\begin{equation*}
\left[P_{m}, P_{n}\right]=m \Delta \delta_{m,-n} \tag{2.26}
\end{equation*}
$$

Note that the Virasoro and Kac-Moody algebra is factorized in this basis. This is presented in this form in [32], which
also gives the relation between this and the algebra presented in [6].

## III. $T \bar{T}$ DEFORMED CSS BOUNDARY CONDITIONS

To compute the $T \bar{T}$ deformed bulk metric corresponding to the $T \bar{T}$ deformed boundary WCFT, we first compute the deformed boundary metric using (2.7):

$$
\begin{align*}
\gamma_{i j}(\lambda) d z^{i} d z^{j}= & -\left(d \bar{z}+\left(\lambda L(z)-P^{\prime}(z)\right) d z\right) \\
& \times\left(d z+\lambda \Delta\left(d \bar{z}-P^{\prime}(z) d z\right)\right) \tag{3.1}
\end{align*}
$$

This metric is flat, so we express it in explicitly flat coordinates with indices $a$ and $b$ :

$$
\begin{equation*}
\gamma_{a b}(\lambda) d u^{a} d u^{b}=-d u d v \tag{3.2}
\end{equation*}
$$

Equating the two, we can calculate the state-dependent coordinate transformation for a $T \bar{T}$ deformed WCFT, analogous to the ones introduced in [33,34]:

$$
\begin{align*}
d u & =d z+\lambda \Delta\left(d \bar{z}-P^{\prime}(z) d z\right) \\
d v & =d \bar{z}+d z\left(\lambda L(z)-P^{\prime}(z)\right) \\
d z & =\frac{d u-\lambda \Delta d v}{1-\lambda^{2} \Delta L(z)} \\
d \bar{z} & =\frac{\left(P^{\prime}(z)-\lambda L(z)\right) d u+\left(\lambda \Delta P^{\prime}(z)-1\right) d v}{1-\lambda^{2} \Delta L(z)} \tag{3.3}
\end{align*}
$$

Furthermore, we can use the flow equations to compute the full bulk metric dual to the $T \bar{T}$ deformed WCFT:

$$
\begin{align*}
\frac{d s^{2}}{l^{2}}= & \frac{d r^{2}}{r^{2}}+\frac{\left(d u\left(\lambda \Delta L+k r^{2}\right)-\Delta d v\left(\lambda k r^{2}+1\right)\right)}{k^{2} r^{2}\left(\lambda^{2} \Delta L-1\right)^{2}} \\
& \times\left(d u\left(\lambda k r^{2}+1\right) L-d v\left(\lambda \Delta L+k r^{2}\right)\right) \tag{3.4}
\end{align*}
$$

where $L \equiv L(u, v)=L(z)$. Note that, on doing so, we lose the $P(z)$ degree of freedom, since this is equivalent to imposing Dirichlet boundary conditions at the constant radial surface $r_{c}=\sqrt{-\frac{1}{k \lambda}}$. If we are to impose DirichletNeumann boundary conditions on this surface, we can recover the $\mathrm{U}(1)$ degree of freedom. To do so, we have to perform the transformation

$$
\begin{equation*}
u \rightarrow u-\lambda \Delta P(u, v), \quad v \rightarrow v-P(u, v) \tag{3.5}
\end{equation*}
$$

This is the analog of the warped symmetry transformation but now in the state-dependent coordinates. We will explore both types of boundary conditions, starting with the simpler case of only imposing Dirichlet boundary conditions.

## A. Asymptotic Killing vectors I: Dirichlet boundary conditions

We will first compute the $T \bar{T}$ deformed asymptotic symmetries which preserve the deformed boundary conditions, which is equivalent to imposing Dirichlet boundary conditions at the radial cutoff surface.

Preserving radial gauge (2.10), we see that the asymptotic Killing vector in the deformed spacetime has the form

$$
\begin{align*}
\xi^{r}(\lambda)= & r f(u, v) \\
\xi^{u}(\lambda)= & V^{u}(u, v)-\frac{k}{k^{2} r^{4}-\Delta L}\left(\Delta\left(2 \lambda k r^{2}+\lambda^{2} \Delta L+1\right) \partial_{u} f\right. \\
& \left.+\left(\lambda \Delta L\left(2+k \lambda r^{2}\right)+k r^{2}\right) \partial_{v} f\right) \\
\xi^{v}(\lambda)= & V^{v}(u, v)-\frac{k}{k^{2} r^{4}-\Delta L}\left(L\left(2 \lambda k r^{2}+\lambda^{2} \Delta L+1\right) \partial_{v} f\right. \\
& \left.+\left(\lambda \Delta L\left(2+k \lambda r^{2}\right)+k r^{2}\right) \partial_{u} f\right) \tag{3.6}
\end{align*}
$$

It will be convenient to define
$W^{u}(u, v)=V^{u}(u, v)+k \lambda \partial_{\bar{z}} f(u, v)$,
$W^{v}(u, v)=V^{v}(u, v)+k \lambda\left(\partial_{z} f(u, v)+P^{\prime}(u, v) \partial_{\bar{z}} f(u, v)\right)$,
where, using (3.3), the derivatives in $z$ and $\bar{z}$ are

$$
\begin{equation*}
\partial_{\bar{z}}=\lambda \Delta \partial_{u}+\partial_{v}, \quad \partial_{z}=\partial_{u}+\lambda L(u, v) \partial_{v}-P^{\prime}(u, v) \partial_{\bar{z}} \tag{3.8}
\end{equation*}
$$

In terms of $W^{a}$, the mixed boundary condition, or, equivalently, the Dirichlet boundary condition at $r=r_{c}$,

$$
\begin{equation*}
\left.\mathcal{L}_{\xi(\lambda)} g_{\mu \nu}(\lambda)\right|_{r=r_{c}}=0, \tag{3.9}
\end{equation*}
$$

constrains the functions in $\xi(\lambda)$ to obey

$$
\begin{align*}
f(u, v) & =-\frac{1}{2}\left(\frac{1-\lambda^{2} \Delta L}{1+\lambda^{2} \Delta L}\right)\left(\partial_{u} W^{u}+\partial_{v} W^{v}\right), \\
W^{u} & =-\left(\frac{\lambda \Delta}{1+\lambda^{2} \Delta L}\right)\left(\partial_{u} W^{u}+\partial_{v} W^{v}\right), \\
W^{v} & =-\left(\frac{\lambda L}{1+\lambda^{2} \Delta L}\right)\left(\partial_{u} W^{u}+\partial_{v} W^{v}\right) . \tag{3.10}
\end{align*}
$$

It turns out that this is not enough to solve for $\delta L$. In the undeformed case (2.15), the functions in the asymptotic Killing vector are all holomorphic functions, so we apply the holomorphicity property in the deformed case as well:

$$
\begin{equation*}
\partial_{\bar{z}} W^{a}(u, v)=0, \quad \partial_{\bar{z}} L(u, v)=0 \tag{3.11}
\end{equation*}
$$

Combining the previous two equations, we get the conditions

$$
\begin{align*}
f(u, v) & =-\frac{1}{2}\left(1-\lambda^{2} \Delta L(u, v)\right) \partial_{u} W^{u}(u, v) \\
\partial_{v} L(u, v) & =-\lambda \Delta \partial_{u} L(u, v) \\
\partial_{v} W^{u}(u, v) & =-\lambda \Delta \partial_{u} W^{u}(u, v) \\
\partial_{a} W^{v}(u, v) & =-\lambda L(u, v) \partial_{a} W^{u}(u, v) \tag{3.12}
\end{align*}
$$

We can use these equations to eliminate $v$ derivatives of all the functions and all derivatives of $W^{v}$.

Now we have enough information to be able to solve for $\delta L$. To do so, we solve

$$
\begin{equation*}
\mathcal{L}_{\xi(\lambda)} g_{\mu \nu}(\lambda)=\partial_{L(u, v)} g_{\mu \nu} \delta_{\xi} L(u, v) \tag{3.13}
\end{equation*}
$$

There are three equations, but, with the relations in (3.12), all three equations become identical, and the $r$ dependence drops out. Solving for $\delta L$, we get

$$
\begin{align*}
\delta_{\xi} L(u, v)= & \left(W^{u}-\lambda \Delta W^{v}\right) L^{\prime} \\
& +\frac{1}{2}\left(4 L+\lambda^{2} k \Delta\left(1-\lambda^{2} \Delta L\right) L^{\prime \prime}\right)\left(1-\lambda^{2} \Delta L\right) W^{u \prime} \\
& +\lambda^{2} k \Delta\left(1-\lambda^{2} \Delta L\right)^{2} L^{\prime} W^{u \prime \prime}-\frac{k}{2}\left(1-\lambda^{2} \Delta L\right)^{3} W^{u \prime \prime \prime} \tag{3.14}
\end{align*}
$$

where ${ }^{\prime}=\partial_{u}$. When $\lambda \rightarrow 0$, we recover (2.17). Note that this depends on two arbitrary functions $W^{u}$ and $W^{v}$.

## 1. Deformed charge algebra

To compute the symmetry algebra of the $T \bar{T}$ deformed holographic WCFT, we must compute the conserved charge algebra of the dual spacetime. We first compute the deformed boundary energy-momentum tensor using the flow equations, which also coincides with the $\mathrm{AdS}_{3}$ BrownYork energy-momentum tensor evaluated on the constant radial surface $r=r_{c}$ :

$$
T_{i j}^{(\lambda)}=-\frac{l}{2 \pi}\left(\begin{array}{cc}
\frac{L}{1-\lambda^{2} \Delta L} & \frac{1+\lambda k+\lambda^{2} \Delta L}{\lambda^{2} k\left(1-\lambda^{2} \Delta L\right)}  \tag{3.15}\\
\frac{1+\lambda k+\lambda^{2} \Delta L}{\lambda^{2} k\left(1-\lambda^{2} \Delta L\right)} & \frac{\Delta}{1-\lambda^{2} \Delta L}
\end{array}\right)
$$

Conserved charges are defined with respect to a constant time coordinate $t$, which is defined in terms of $u$ and $v$ by

$$
\begin{equation*}
u=t+\phi, \quad v=t-\phi \tag{3.16}
\end{equation*}
$$

Since $L$ is holomorphic in $z$, we can express the $t$ derivative in terms of the $\phi$ derivative:

$$
\begin{equation*}
\partial_{t} L=\frac{1+\lambda \Delta}{1-\lambda \Delta} \partial_{\phi} L \tag{3.17}
\end{equation*}
$$

So we can express the $u$ derivatives of holomorphic functions only in $\phi$ derivatives as well:

$$
\begin{equation*}
\partial_{u}=\frac{1}{2}\left(\partial_{t}+\partial_{\phi}\right)=\frac{1}{1-\lambda \Delta} \partial_{\phi} \tag{3.18}
\end{equation*}
$$

For constant $t$, we can now eliminate $W^{v}$ in (3.14), using Eqs. (3.12) and (3.18):
$\partial_{u} W^{v}=-\lambda L(\phi) \partial_{u} W^{u} \Rightarrow \partial_{\phi} W^{v}=-\lambda L(\phi) \partial_{\phi} W^{u}$.
Integrating over $\phi$, we have

$$
\begin{equation*}
W^{v}(\phi)=\int^{\phi} d \phi^{\prime} L\left(\phi^{\prime}\right) \partial_{\phi^{\prime}} W^{u}\left(\phi^{\prime}\right) \tag{3.20}
\end{equation*}
$$

Now we can label the variation of the conserved charges with only one arbitrary function $W^{u}$. Using (2.21) but with the deformed energy-momentum tensor, the conserved charge is

$$
\begin{equation*}
Q_{f}=\frac{l}{4 \pi^{2}} \int_{0}^{2 \pi} d \phi f(\phi) \frac{\Delta-L(\phi)}{1-\lambda^{2} \Delta L(\phi)} \tag{3.21}
\end{equation*}
$$

We can now compute the charge algebra:

$$
\begin{align*}
\left\{Q_{W}, Q_{f}\right\}= & \delta_{W} Q_{f} \\
= & \frac{l}{4 \pi^{2}} \int_{0}^{2 \pi} d \phi f \\
& \times\left(\frac{-\delta_{W} L}{1-\lambda^{2} \Delta L}+\frac{\Delta-L}{\left(1-\lambda^{2} \Delta L\right)^{2}}\left(\lambda^{2} \Delta \delta_{W} L\right)\right) \tag{3.22}
\end{align*}
$$

Using (3.14) and (3.20), substituting $f(\phi)=e^{i m \phi}$ and $W(\phi)=e^{i n \phi}$, and removing $\phi$ derivatives from $L$ using integration by parts, we have

$$
\begin{align*}
\left\{Q_{W}, Q_{f}\right\}= & \delta_{W} Q_{f}=\frac{l(1+\lambda \Delta)}{8 \pi^{2}(1-\lambda \Delta)^{2}} \int_{0}^{2 \pi} d \phi \frac{1}{1-\lambda^{2} \Delta L(\phi)}\left[2 i n^{3} k(1-\lambda \Delta)^{3} e^{i(m+n) \phi}\left(1-\lambda^{2} \Delta L(\phi)\right)\right. \\
& +e^{i m \phi} L(\phi)\left(2 m n \lambda \Delta(1-\lambda \Delta)^{2} \int^{\phi} e^{i n \phi^{\prime}} L\left(\phi^{\prime}\right) d \phi^{\prime}-i e^{i n \phi}\left(n \lambda^{2} k \Delta\left(1-\lambda^{2} \Delta L(\phi)\right)\right.\right. \\
& \left.\left.\left.\times\left(n^{2} \lambda \Delta(3-\lambda \Delta(3-\lambda \Delta))-m^{2}\right)\right)-2(1-\lambda \Delta)^{2}((m-n)-2 n \lambda \Delta L(\phi))\right)\right] . \tag{3.23}
\end{align*}
$$

Since $L$ is not independent of $t$, only the zero modes are conserved in time. In this choice of basis of functions and Fourier modes, the central charge term is state dependent. This is similar to what was found in [23] for a $T \bar{T}$ deformed CFT. It is straightforward to verify that, on taking the $\lambda \rightarrow 0$ limit and expressing $L$ in Fourier modes, one recovers the Virasoro algebra.

## B. Asymptotic Killing vectors II: Dirichlet-Neumann boundary conditions

If we want to impose the same boundary conditions at the radial cutoff in the $T \bar{T}$ deformed metric as the undeformed Dirichlet-Neumann CSS boundary conditions at infinity of the undeformed metric, we have to find a global Killing vector which corresponds to translations on the boundary. It is easy to verify that $\lambda \Delta \partial_{u}+\partial_{v}$ is such a global Killing vector of (3.4). To generate transformations in the boundary metric, we then promote this global Killing vector to an "asymptotic Killing vector" analogous to (2.18):

$$
\begin{equation*}
\eta(\lambda ; \sigma)=-\sigma(u, v)\left(\lambda \Delta \partial_{u}+\partial_{v}\right) . \tag{3.24}
\end{equation*}
$$

To introduce the $P(z)$ degree of freedom back into the metric (3.4), one can make the coordinate transformation (3.5):

$$
\begin{equation*}
u \rightarrow u-\lambda \Delta h(u, v), \quad v \rightarrow v-h(u, v), \tag{3.25}
\end{equation*}
$$

which is generated by the asymptotic Killing vector (3.24) as the analog to the warped transformation $\bar{z} \rightarrow \bar{z}-h(z)$. The state-dependent coordinate transformation is now

$$
\begin{align*}
d u-\lambda \Delta \mathrm{d}(h(u, v))= & d z+\lambda \Delta \mathrm{d}(\bar{z}-P(z)), \\
d v-\mathrm{d}(h(u, v))= & d \bar{z}+\left(\lambda L(z)-P^{\prime}(z)\right) d z, \\
d z= & \frac{d u-\lambda \Delta d v}{1-\lambda^{2} \Delta L}, \\
d \bar{z}= & \frac{d v-\lambda L d u+(d u+\lambda \Delta d v) P^{\prime}(z)}{1-\lambda^{2} \Delta L} \\
& -(\mathrm{d}(h(u, v)), \tag{3.26}
\end{align*}
$$

where d is the exterior derivative. Note that, since both $h$ and $P$ are arbitrary functions of $(u, v)$, we can choose the gauge where $h=P$. The coordinate transformation now becomes much simpler:

$$
\begin{array}{lr}
d u=d z+\lambda \Delta \bar{z}, & d v=d \bar{z}+\lambda L(z) d z, \\
d z=\frac{d u-\lambda \Delta d v}{1-\lambda^{2} \Delta L}, & d \bar{z}=\frac{d v-\lambda L d u}{1-\lambda^{2} \Delta L} . \tag{3.27}
\end{array}
$$

The metric now reads

$$
\begin{align*}
d s^{2}= & l^{2} \frac{d r^{2}}{r^{2}}+\frac{l^{2}}{k^{2} r^{2}\left(1-\lambda^{2} \Delta L\right)^{2}}\left(\left(k r^{2}\left(\lambda^{2} \Delta L-1\right) \partial_{u} h-\left(1+\lambda k r^{2}\right) L\right) d u+\left(k r^{2} \partial_{v} h\left(\lambda^{2} \Delta L-1\right)+\lambda \Delta L+k r^{2}\right) d v\right) \\
& \times\left(\left(\Delta \partial_{u} h\left(\lambda^{2} \Delta L-1\right)-\lambda \Delta L-k r^{2}\right) d u+\Delta\left(\partial_{v} h\left(\lambda^{2} \Delta L-1\right)+\lambda k r^{2}+1\right) d v\right) . \tag{3.28}
\end{align*}
$$

This metric still has the asymptotic Killing vector (3.24), and when $\sigma=1$ it is a global Killing vector. Computing the flow in phase space generated by (3.24), we have

$$
\begin{equation*}
\mathcal{L}_{\eta(\lambda, \sigma)} g_{\mu \nu}\left(\lambda ; L(u, v), \partial_{u} h(u, v)\right)=\partial_{L} g_{\mu \nu} \delta L+\partial_{h} g_{\mu \nu} \delta h, \tag{3.29}
\end{equation*}
$$

which, on solving, we see that we recover the undeformed $U(1)$ symmetry:

$$
\begin{equation*}
\delta L=0, \quad \delta h=\sigma . \tag{3.30}
\end{equation*}
$$

Now we will see if on performing the warped transformation (3.25) we lose the deformed Virasoro symmetry (3.14). The vector field which preserves radial gauge is

$$
\begin{align*}
\xi^{r}= & r f(u, v), \\
\xi^{u}= & V^{u}(u, v)+\frac{k}{\left(\Delta L-k^{2} r^{4}\right)\left(1-\lambda \Delta \partial_{u} h-\partial_{v} h\right)^{2}}\left[\partial _ { v } f \left(\Delta \lambda L \left(\Delta \lambda \partial_{u} h\left(-2 \partial_{v} h+k \lambda r^{2}+1\right)\right.\right.\right. \\
& \left.\left.+\left(1-\lambda k r^{2}\right) \partial_{v} h-\lambda k r^{2}-2\right)+\Delta \partial_{u} h\left(\partial_{v} h-\lambda k r^{2}-1\right)+k r^{2}\left(\partial_{v} h-1\right)+\lambda^{3} \Delta^{2} L^{2} \partial_{v} h\left(\lambda \Delta \partial_{u} h-1\right)\right) \\
& \left.-\Delta \partial_{u} f\left(2\left(\lambda k r^{2}+1\right)\left(\lambda^{2} \Delta L-1\right) \partial_{v} h+\left(\partial_{v} h\right)^{2}\left(\lambda^{2} \Delta L-1\right)^{2}+\lambda^{2} \Delta L+2 \lambda k r^{2}+1\right)\right], \\
\xi^{v}= & V^{v}(u, v)+\frac{k}{\left(\Delta L-k^{2} r^{4}\right)\left(1-\lambda \Delta \partial_{u} h-\partial_{v} h\right)^{2}}\left[\partial _ { u } f \left(-\partial_{v} h\left(\Delta \lambda^{2} L-1\right)\left(\Delta \lambda L+k r^{2}\right)\right.\right. \\
& \left.+\Delta \partial_{u} h\left(\Delta \lambda^{2} L-1\right)\left(\partial_{v} h\left(\Delta \lambda^{2} L-1\right)+k \lambda r^{2}+1\right)-\lambda \Delta L\left(\lambda k r^{2}+2\right)-k r^{2}\right) \\
& \left.+\partial_{v} f\left(2 \partial_{u} h\left(\lambda^{2} \Delta L-1\right)\left(\Delta \lambda L+k r^{2}\right)-\Delta\left(\partial_{u} h\right)^{2}\left(\Delta \lambda^{2} L-1\right)^{2}-L\left(\Delta \lambda^{2} L+2 k \lambda r^{2}+1\right)\right)\right] . \tag{3.31}
\end{align*}
$$

To have solutions which preserve the mixed boundary conditions, we are required to impose holomorphicity of $h$ in $z, \bar{z}$ coordinates:

$$
\begin{equation*}
\partial_{\bar{z}} h=0 \Rightarrow \partial_{v} h=-\lambda \Delta \partial_{u} h . \tag{3.32}
\end{equation*}
$$

To simplify the equations, one can introduce the following definitions:

$$
\begin{align*}
W^{u} & =V^{u}-\lambda^{3} k \Delta^{2}\left(1-\lambda^{2} \Delta L\right) \partial_{u} f \partial_{u} h \\
W^{v} & =V^{v}+\lambda k\left(1-\lambda^{2} \Delta L\right) \partial_{u} f\left(1-\lambda \Delta \partial_{u} h\right) \\
X & =W^{u}-\lambda \Delta W^{v} \tag{3.33}
\end{align*}
$$

To compute the variation of the functions $L$ and $h$, it will be necessary to impose holomorphicity in $z$ for the functions $W^{a}, X$, and $f$. We can now compute the variation of the metric which preserves the mixed boundary conditions or, equivalently, impose Dirichlet boundary conditions at the constant radial surface:

$$
\begin{equation*}
\left.\mathcal{L}_{\xi(\lambda, \sigma)} g_{\mu \nu}\left(\lambda ; L(u, v), \quad \partial_{u} h(u, v)\right)\right|_{r=r_{c}}=0 \tag{3.34}
\end{equation*}
$$

This is a set of three equations; however, only two are linearly independent. The conditions we get from solving the above equations are
$f(u, v)=-\frac{1}{2}\left(\frac{1-\lambda^{2} \Delta L}{1+\lambda^{2} \Delta L}\right) X^{\prime}(u, v)$,
$X(u, v)=\frac{W^{v}(u, v)}{h^{\prime \prime}(u, v)}-\frac{h^{\prime}(u, v)-\lambda L\left(1-\lambda \Delta h^{\prime}(u, v)\right) X^{\prime}(u, v)}{h^{\prime \prime}(u, v)\left(1+\lambda^{2} \Delta L\right)}$,
where ${ }^{\prime} \equiv \partial_{u}$.
We are now in a position to compute the flow of the metric in phase space generated by this vector field subject to the above constraints:

$$
\begin{equation*}
\mathcal{L}_{\xi(\lambda, \sigma)} g_{\mu \nu}\left(\lambda ; L(u, v), \partial_{u} h(u, v)\right)=\partial_{L} g_{\mu \nu} \delta L_{\xi}+\partial_{h} g_{\mu \nu} \delta h_{\xi} . \tag{3.36}
\end{equation*}
$$

As before, this set of three equations subject to the constraints reduces to two equations and removes any dependence on the radial coordinate $r$. The variations of the functions $h$ and $L$ are

$$
\begin{align*}
\delta_{\xi} h= & 0 \\
\delta_{\xi} L= & \frac{1}{2 \Theta^{2} h^{\prime \prime}}\left(\left(2 \lambda k L^{\prime} L_{m}^{2} L_{p}\left(2 \lambda \Delta h^{\prime}-1\right) h^{\prime \prime}+3 k L_{m}^{3} L_{p}^{2}\left(h^{\prime \prime}\right)^{2}\right) W^{\prime \prime}-k \Theta L_{m}^{3} L_{p} h^{\prime \prime} W^{\prime \prime \prime}\right. \\
& +X^{\prime}\left(-2 \lambda^{2} k\left(L^{\prime}\right)^{2} L_{m}^{2} h^{\prime \prime}\left(1-3 \lambda \Delta h^{\prime}+2 \lambda^{2} \Delta^{2}\left(h^{\prime}\right)^{2}\right)-6 k L_{m}^{3} L_{p}^{2}\left(h^{\prime \prime}\right)^{3}+h^{\prime \prime}\left(-\Theta L_{m}\left(\lambda k L^{\prime \prime}\left(1-2 \lambda \Delta h^{\prime}\right)\right.\right.\right. \\
& \left.\left.+4 \lambda L^{2}\left(1-\lambda \Delta h^{\prime}\right)+L\left(-\lambda^{3} k \Delta L^{\prime \prime}-2\left(2-\lambda^{4} k \Delta^{2} L^{\prime \prime}\right) h^{\prime}\right)\right)+6 k \Theta L_{m}^{3} L_{p} h^{\prime \prime \prime}\right)-L^{\prime}\left(2 \Theta^{3}\right. \\
& \left.\left.+\lambda k L_{m}^{2}\left(-7+8 \lambda \Delta h^{\prime}-\lambda^{2} \Delta L\left(1-8 \lambda \Delta h^{\prime}\right)\right)\left(h^{\prime \prime}\right)^{2}+2 \lambda k \Theta L_{m}^{2}\left(1-2 \lambda \Delta h^{\prime}\right) h^{\prime \prime \prime}\right)-k \Theta^{2} L_{m}^{3} h^{\prime \prime \prime \prime}\right) \\
& \left.+W^{\prime}\left(L^{\prime}\left(2 \Theta^{2} L_{p}+2 \lambda k L_{m}^{2} L_{p}\left(1-2 \lambda \Delta h^{\prime}\right) h^{\prime \prime \prime}+k \Theta L_{m}^{3} L_{p} h^{\prime \prime \prime \prime}\right)-3 k L_{m}^{3} L_{p}^{2} h^{\prime \prime} h^{\prime \prime \prime}\right)\right), \tag{3.37}
\end{align*}
$$

where

$$
\begin{align*}
\Theta & =\partial_{u} h-\lambda L\left(1-\lambda \Delta \partial_{u} h\right), \quad W=W^{v}, \\
L_{m} & =1-\lambda^{2} \Delta L, \quad L_{p}=1+\lambda^{2} \Delta L . \tag{3.38}
\end{align*}
$$

We see that we still preserve a deformed Virasoro generator and do not generate a transformation in the $\mathrm{U}(1)$ generator. However, the algebra produced the modes of the charges will not be closed, as the variation of $L$ depends on the $\mathrm{U}(1)$ generator $h$. We will not compute a charge algebra for this, since it is not illuminating but, in principle, can be computed using the same procedure outlined in the previous section.

Let us compare the results of this section with Sec. III A. We find that the $T \bar{T}$ deformation does not affect the spin 1 currents, and, therefore, the deformed theory should retain whatever Kac-Moody algebra the undeformed theory has. This suggests that the results in Sec. III A are only a special case of this section, where the bulk dual is dual to a state with zero momentum in the boundary deformed WCFT.

## IV. DISCUSSION

In this paper, we computed the $T \bar{T}$ deformed generators of a warped CFT, using holographic techniques developed in [23,27]. Previously, holographic $T \bar{T}$ techniques have been used to compute $T \bar{T}$ deformations of holographic CFTs. Since the $T \bar{T}$ deformation is a double-trace deformation, the boundary conditions of the holographic bulk dual are modified. For the $T \bar{T}$ deformation, this can be interpreted as imposing Dirichlet boundary conditions at a finite radial surface for the bulk metric. However, when considering a holographic WCFT dual to $\mathrm{AdS}_{3}$, one has to employ the CSS boundary conditions [6], which are Dirichlet-Neumann boundary conditions for the bulk metric.

We, therefore, computed the $T \bar{T}$ deformed CSS boundary conditions, by imposing either only Dirichlet boundary conditions at the cutoff radial surface or Dirichlet-Neumann boundary conditions at the same surface. Using this, we computed the $T \bar{T}$ deformed asymptotic symmetry algebra for both cases and found that, for a $T \bar{T}$
deformed holographic WCFT, the U(1) Kac-Moody generators are not affected, but the Virasoro generators are deformed in a nonlinear way. In fact, when considering the Dirichlet-Neumann boundary conditions at the finite radial surface, we see that the deformed Virasoro generator will no longer create a closed algebra with itself, but the full deformed asymptotic algebra is still closed. This suggests that the symmetry algebra of the $T \bar{T}$ deformed WCFT still contains the $\mathrm{U}(1)$ Kac-Moody algebra, which follows from the fact that the $T \bar{T}$ deformation preserves translation invariance.

A natural question to ask now is which of the two boundary conditions corresponds to the correct $T \bar{T}$ flow of the boundary field theory. Since the $T \bar{T}$ deformation does not effect spin 1 currents, the deformed theory should not lose the $\mathrm{U}(1)$ Kac-Moody algebra, which suggests that the second approach yields the correct deformed theory.

This result strengthens and extends the proposals of [22,23,27] to the case of an example of bottom-up holography where the boundary theory is not a conformal field theory but instead a nonrelativistic theory. It will be interesting to explore how the holographic $T \overline{\bar{T}}$ dictionary extends to other examples of holography and, in particular, non-AdS holography.

There are many directions one can take from here. Another starting point for a bulk dual to a nonrelativistic QFT would be a $J \bar{T}$ deformation of a CFT dual to $\mathrm{AdS}_{3}$ with a U(1) Chern-Simons matter field, to generate CSSlike boundary conditions [35]. Since warped CFTs can also be formulated as dual to modified gravity theories with a warped AdS bulk, it would also be interesting to use the Chern-Simons formalism of holographic $T \bar{T}$ [30] to compute the $T \bar{T}$ deformations of WCFT dual to warped $\mathrm{AdS}_{3}$ as a solution to lower spin gravity [3] or as a solution to massive gravity [36,37]. Stepping away from holography, it would be interesting to compute the $T \bar{T}$ deformed WCFT partition function and explore the deformations of other nonrelativistic QFTs such as the quantum Lifshitz model in $2+1 \mathrm{D}$, which will require understanding $T \bar{T}$ deformations in higher dimensions.

## ACKNOWLEDGMENTS

The author thanks Valentina Giangreco M. Puletti, Monica Guica, and Lárus Thorlacius for their insights and helpful discussions. This work was supported by the Icelandic Research Fund under Grant No. 228952-052 and by a University of Iceland doctoral grant.
[1] D. M. Hofman and A. Strominger, Chiral scale and conformal invariance in 2D quantum field theory, Phys. Rev. Lett. 107, 161601 (2011).
[2] S. Detournay, T. Hartman, and D. M. Hofman, Warped conformal field theory, Phys. Rev. D 86, 124018 (2012).
[3] D. M. Hofman and B. Rollier, Warped conformal field theory as lower spin gravity, Nucl. Phys. B897, 1 (2015).
[4] J. M. Bardeen and G. T. Horowitz, The extreme Kerr throat geometry: A vacuum analog of $\mathrm{AdS}_{2} \times S^{2}$, Phys. Rev. D 60, 104030 (1999).
[5] O. J. C. Dias, R. Emparan, and A. Maccarrone, Microscopic theory of black hole superradiance, Phys. Rev. D 77, 064018 (2008).
[6] G. Compère, W. Song, and A. Strominger, New boundary conditions for $\mathrm{AdS}_{3}$, J. High Energy Phys. 05 (2013) 152.
[7] D. Anninos, J. Samani, and E. Shaghoulian, Warped entanglement entropy, J. High Energy Phys. 02 (2014) 118.
[8] A. Castro, D. M. Hofman, and N. Iqbal, Entanglement entropy in warped conformal field theories, J. High Energy Phys. 02 (2016) 033.
[9] W. Song, Q. Wen, and J. Xu, Modifications to holographic entanglement entropy in warped CFT, J. High Energy Phys. 02 (2017) 067.
[10] L. Apolo, H. Jiang, W. Song, and Y. Zhong, Modular Hamiltonians in flat holography and (W)AdS/WCFT, J. High Energy Phys. 09 (2020) 033.
[11] A. Castro, C. Keeler, and P. Szepietowski, Tweaking one-loop determinants in $\mathrm{AdS}_{3}$, J. High Energy Phys. 10 (2017) 070.
[12] A. B. Zamolodchikov, Expectation value of composite field T anti-T in two-dimensional quantum field theory, arXiv: hep-th/0401146.
[13] F. A. Smirnov and A. B. Zamolodchikov, On space of integrable quantum field theories, Nucl. Phys. B915, 363 (2017).
[14] M. Caselle, D. Fioravanti, F. Gliozzi, and R. Tateo, Quantisation of the effective string with TBA, J. High Energy Phys. 07 (2013) 071.
[15] A. Cavaglià, S. Negro, I. M. Szécsényi, and R. Tateo, $T \bar{T}$ deformed 2D quantum field theories, J. High Energy Phys. 10 (2016) 112.
[16] S. Dubovsky, V. Gorbenko, and M. Mirbabayi, Natural tuning: Towards a proof of concept, J. High Energy Phys. 09 (2013) 045.
[17] S. Dubovsky, V. Gorbenko, and G. Hernández-Chifflet, $T \bar{T}$ partition function from topological gravity, J. High Energy Phys. 09 (2018) 158.
[18] J. Cardy, The $T \bar{T}$ deformation of quantum field theory as random geometry, J. High Energy Phys. 10 (2018) 186.
[19] J. Cardy, $T \bar{T}$ deformation of correlation functions, J. High Energy Phys. 12 (2019) 160.
[20] O. Aharony and N. Barel, Correlation functions in $T \bar{T}$ deformed conformal field theories, J. High Energy Phys. 08 (2023) 035.
[21] N. Benjamin, S. Collier, J. Kruthoff, H. Verlinde, and M. Zhang, S-duality in $T \bar{T}$-deformed CFT, J. High Energy Phys. 05 (2023) 140.
[22] L. McGough, M. Mezei, and H. Verlinde, Moving the CFT into the bulk with $T \bar{T}$, J. High Energy Phys. 04 (2018) 010.
[23] M. Guica and R. Monten, $T \bar{T}$ and the mirage of a bulk cutoff, SciPost Phys. 10, 024 (2021).
[24] I. R. Klebanov and E. Witten, AdS/CFT correspondence and symmetry breaking, Nucl. Phys. B556, 89 (1999).
[25] E. Witten, Multitrace operators, boundary conditions, and AdS/CFT correspondence, arXiv:hep-th/0112258.
[26] L. Apolo, P.-X. Hao, W.-X. Lai, and W. Song, Glue-on AdS holography for $T \bar{T}$-deformed CFTs, J. High Energy Phys. 06 (2023) 117.
[27] P. Kraus, R. Monten, and R. M. Myers, 3D gravity in a box, SciPost Phys. 11, 070 (2021).
[28] M. He, S. He, and Y.-h. Gao, Surface charges in ChernSimons gravity with $T \bar{T}$ deformation, J. High Energy Phys. 03 (2022) 044.
[29] K. Skenderis and S. N. Solodukhin, Quantum effective action from the AdS/CFT correspondence, Phys. Lett. B 472, 316 (2000).
[30] E. Llabrés, General solutions in Chern-Simons gravity and $T \bar{T}$-deformations, J. High Energy Phys. 01 (2021) 039.
[31] M. Banados, C. Teitelboim, and J. Zanelli, The black hole in three-dimensional space-time, Phys. Rev. Lett. 69, 1849 (1992).
[32] L. Apolo and W. Song, Bootstrapping holographic warped CFTs or: How I learned to stop worrying and tolerate negative norms, J. High Energy Phys. 07 (2018) 112.
[33] S. Dubovsky, R. Flauger, and V. Gorbenko, Solving the simplest theory of quantum gravity, J. High Energy Phys. 09 (2012) 133.
[34] S. Dubovsky, V. Gorbenko, and M. Mirbabayi, Asymptotic fragility, near $\mathrm{ads}_{2}$ holography and $t \bar{T}$, J. High Energy Phys. 09 (2017) 136.
[35] A. Bzowski and M. Guica, The holographic interpretation of $J \bar{T}$-deformed CFTs, J. High Energy Phys. 01 (2019) 198.
[36] S. Deser, R. Jackiw, and S. Templeton, Topologically massive gauge theories, Ann. Phys. (N.Y.) 140, 372 (1982).
[37] E. A. Bergshoeff, O. Hohm, and P. K. Townsend, Massive gravity in three dimensions, Phys. Rev. Lett. 102, 201301 (2009).


[^0]:    *rap19@hi.is

[^1]:    ${ }^{2}$ The quotes are to indicate that, since this vector does not satisfy Dirichlet boundary conditions, it is technically not an asymptotic Killing vector, but, since it generates flows in the phase space, it will continue to be referred to as such later in this paper.

