Type III seesaw with R-parity violation in light of m_W (CDF)

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Abstract

Motivated by the recently reported measurement of the W boson mass $M_W = 80.4335 \pm 0.0094$ GeV by the CDF collaboration, we propose a type III seesaw extension of the minimal supersymmetric standard model (MSSM) which also includes an R-parity violating term. Without taking potential SUSY radiative corrections into account, we show that the CDF measurement of M_W and the LEP measurement of the ρ parameter can be simultaneously accommodated at the 2σ level. A long-lived gravitino in a few GeV mass range is a unique viable dark matter candidate in this framework.

1. Introduction

The CDF collaboration at Fermilab has reported a very precise measurement of the W boson pole mass $m_W^{\mathrm{CDF}} = 80.4335 \pm 0.0094$ GeV [1], which shows about 7σ deviation from the

Standard Model (SM) prediction, $m_W^{\text{SM}} = 80.357 \pm 0.006$ GeV [2]. For the central values, the deviation is $\Delta m_W = m_W^{\text{CDF}} - m_W^{\text{SM}} = 0.0765$ GeV, which may be viewed as tantalizing evidence for the new physics beyond the SM. Following the CDF report, several works have examined new physics scenarios to account for the observed deviation [3–33].

In addition to the CDF results, one should also consider the electroweak precision measurements (EWPM) by the LEP experiment. With the very precisely measured Z boson mass m_Z by the LEP experiment [34], the SM prediction for the ρ-parameter is consistent with the LEP result, $\rho \equiv \frac{m_W^2}{m_Z^2 c_W^2} = 1.0003 \pm 0.0005$ [35]. Hence, a large deviation of W boson mass from its SM predicted value is likely in tension with the LEP result. The challenge for any new physics scenario is to minimize the tension between the two results.

In this letter, we provide a simple extension of the MSSM which incorporates type III seesaw and which can accommodate to a large extent the Δm_W in the presence of a suitable R-parity violating term. The model is also consistent with the LEP results and satisfies various other phenomenological constraints, in addition to providing a framework for neutrino masses and mixings. With the inclusion of the gauge-mediated supersymmetry breaking, the model can also accommodate successful inflationary scenario. Finally, the model can be embedded within grand unified theory (GUT) such as $SU(5)$.

2. The model

MSSM is a well-motivated candidate for physics beyond the SM that provides a natural resolution for the SM gauge hierarchy problem [36–40]. However, it cannot explain the origin of the observed neutrino mass as shown by the solar and atmospheric neutrino oscillation experiments [34]. To remedy this, we implement type III seesaw by including two $SU(2)_L$ triplet chiral 'matter' superfields, Δ_i ($i = 1, 2$) with zero hypercharge. The MSSM superpotential is extended to include

$$W \supset \sum_{i=1}^{2} \sum_{j=1}^{3} \sqrt{2} Y_D^{ij} H_u^T \varepsilon \Delta_i L_j - \frac{m_\Delta}{2} \sum_{i=1}^{2} Tr\left[\Delta_i \Delta_i\right]. \tag{1}$$

Here H_u and L_i respectively denote the up-type Higgs doublet and the lepton doublet and

$$\varepsilon = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}, \qquad \Delta_i = \frac{1}{\sqrt{2}} \begin{pmatrix} \Delta_i^0 & \sqrt{2}\Delta_i^+ \\ \sqrt{2}\Delta_i^- & -\Delta_i^0 \end{pmatrix}. \tag{2}$$

The Higgs doublets, $H_{u,d}$ develop nonzero vacuum expectation value (VEVs) in the usual way, which break the electroweak symmetry, $\langle H_u \rangle = \left(0, v_u/\sqrt{2}\right)^T$ and $\langle H_d \rangle = \left(v_d/\sqrt{2}, 0\right)^T$, where $v_u = v \cos\beta$, $v_d = v \sin\beta$ and $v = \sqrt{v_u^2 + v_d^2} = 246$ GeV, and H_d is the down-type Higgs doublet. Assuming $m_\Delta \gg Y_D v_u$, the light Majorana neutrino mass matrix generated via the type III seesaw mechanism shown in Fig. 1 (a) is given by

$$m_\nu = \frac{Y_D^T Y_D v_u^2}{2m_\Delta} \sim \frac{Y_D^T Y_D v^2}{2m_\Delta}, \tag{3}$$

for $\tan\beta = v_u/v_d \gtrsim 10$.

We also introduce the R-parity (lepton-number) violating term in the superpotential,

$$W_{\not{R}} = \sqrt{2}\lambda H_u^T \varepsilon \Delta_1 H_d. \tag{4}$$

Note that we have only used Δ_1 to break R-parity for simplicity. The inclusion of this term is crucial to generate an induced VEV for $\widetilde{\Delta}_1$ (the scalar component of Δ_1). This helps increase the W boson mass at tree-level, without altering the Z boson mass, which is consistent with the LEP result. To illustrate this, we consider the trilinear scalar soft SUSY breaking term corresponding to $W_{\not{R}}$ which is of the following form:

$$V \supset \sqrt{2}\lambda A H_u^T \varepsilon \widetilde{\Delta} H_d + h.c., \tag{5}$$

where the parameter A has mass dimension of one. After the electroweak symmetry breaking, and taking into account the mass squared term for $\widetilde{\Delta}_1^0$ in Eq. (1), the relevant potential is given by

$$V \supset m_\Delta^2 |\widetilde{\Delta}_1^0|^2 - \frac{1}{2}\lambda A v_u v_d \widetilde{\Delta}_1^0 + h.c.. \tag{6}$$

As a result $\widetilde{\Delta}_1^0$ acquires an induced VEV,

$$\left\langle \widetilde{\Delta}_1^0 \right\rangle = \frac{1}{2}\lambda A v_u v_d / m_\Delta^2 \equiv v_\Delta / \sqrt{2}. \tag{7}$$

Here, we have neglected the soft SUSY breaking mass-squared term for $\widetilde{\Delta}_1^0$ by assuming it to be much smaller than m_Δ^2. With $v_\Delta \neq 0$, the W and Z masses are given by

$$m_W^2 = \frac{g^2}{4}\left(v^2 + 4v_\Delta^2\right), \tag{8}$$

$$m_Z^2 = \frac{g^2}{4\cos^2\theta_W}v^2, \tag{9}$$

where $\cos\theta_W$ is the electroweak mixing angle. Clearly, since Δ_1 has zero hypercharge, it only contributes to the W boson mass. For $v_\Delta^2 \ll v^2$,

$$\Delta m_W = m_W - m_W^{SM} \simeq 2m_W^{SM}\left(\frac{v_\Delta^2}{v^2}\right). \tag{10}$$

To reproduce the CDF measurement within $n - \sigma$ of the central value, we require

$$v_\Delta^{CDF}(\text{GeV}) = 19.4 \times \sqrt{0.0765 - n \times 0.0094}. \tag{11}$$

As previously discussed, we should also consider the LEP results for EWPM, in particular the ρ parameter. At tree-level[1] it is given by

$$\rho \equiv \frac{m_W^2}{m_Z^2 c_W^2} = 1 + 4\left(\frac{v_\Delta^2}{v^2}\right) \equiv 1 + \Delta\rho. \tag{12}$$

To reproduce the LEP measurement $\rho = 1.0004 \pm 0.0005$ [34,43] within $n - \sigma$ of the central value requires,

$$v_\Delta^{LEP}(\text{GeV}) = 123.4 \times \sqrt{0.0004 + n \times 0.0005}. \tag{13}$$

Clearly, the CDF and LEP results are in tension with each other, which is indicated by the fact that Eqs. (13) and (11) are inconsistent with each other for the central values.
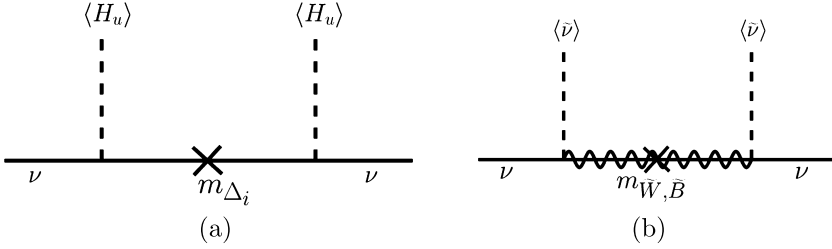
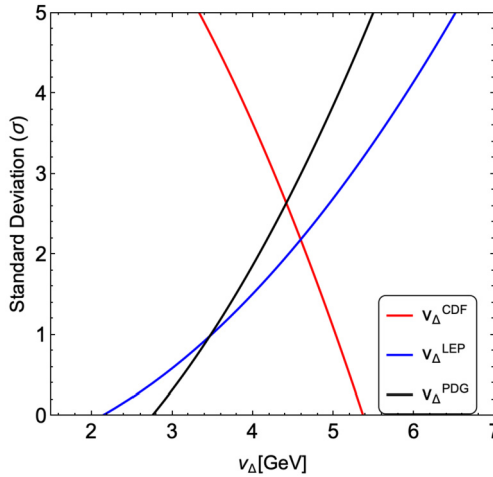Fig. 1. Neutrino mass generation mechanism in the model.

Fig. 2. Induced VEV v_Δ^{CDF} (red), v_Δ^{LEP} (blue), and v_Δ^{PDG} (blue) versus n$-\sigma$ deviation.

In Fig. 2, we show a plot of v_Δ^{CDF} (GeV) and v_Δ^{LEP} (GeV) versus the number of standard deviations n. One needs at least 2σ deviations for both m_W^{CDF} and $\Delta\rho$ to be consistent with each other. In addition to LEP and CDF-II results, we also show a line (black) corresponding to the PDG world average of $M_W^{\text{PDG}} = 80.377 \pm 0.012$ GeV [44], which takes into account the W mass measurements from LEP [45], Tevatron [46] (CDF [47] and D0 [48]) and LHCb collaboration [49].

We therefore conclude that in the type III seesaw extension of MSSM, if an R-parity violating term is introduced, the precisely determined m_W^{CDF} can be accommodated in the model with $v_\Delta \sim 4$ GeV, which is induced by the appearance of the soft-SUSY breaking trilinear scalar coupling term involving the type III seesaw messenger Δ_i.

3. Constraints on R-parity violation

To check the consistency of the model, let us next consider the phenomenological constraints on R-parity violation. We first note that v_Δ generates the so-called bi-linear R-parity violating term:

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[1] References [41,42] include a detailed analysis of oblique correction and global fit to EWPM in the presence of a triplet scalar field. These corrections are negligible compared to this mass shift at the tree-level.

$$W \supset \sqrt{2} \, (Y_b)^{1j} \, H_u^T \varepsilon \langle \widetilde{\Delta}_1 \rangle L_j \equiv \mu_\Delta^j H_u^T \sigma^1 L_j \,, \tag{14}$$

where σ is the Pauli matrix. Together with the MSSM μ-term,

$$W \supset \mu_H H_u^T \varepsilon H_d \,, \tag{15}$$

the H_d and L_j fields mix, and the mixing angles are

$$\varepsilon_{\hat{j}} \sim \frac{\mu_\Delta^j}{\mu_H} \sim (Y_D)^{1j} \frac{v_\Delta}{\mu_H} \,. \tag{16}$$

This leads to R-parity violating terms (lepton number violating Yukawa interactions):

$$W_{\not{R}} \supset Y_e^{ij} E_i^c L_j \left(\sum_k \varepsilon_k L_k \right) + Y_d^{ij} D_i^c Q_j \left(\sum_k \varepsilon_k L_k \right) \tag{17}$$

Such lepton number violating processes can be active in the early universe and wash out the baryon asymmetry. To avoid such a wash-out, we must ensure that these lepton number changing processes fall out of equilibrium for temperature $T \gtrsim \tilde{m}$, where $\tilde{m}$ is a typical sparticle mass. For $\tilde{m} = \mathcal{O}(1\,\text{TeV})$, this requires (see Refs. [50–53])

$$\left(Y_{e,d}\right)^{ij} \varepsilon_k \lesssim 10^{-7}. \tag{18}$$

For the Yukawa couplings to be in perturbative regime, the above conditions is always satisfied if

$$\varepsilon_k \sim \frac{(Y_D)^{1k} \, v_\Delta}{\mu_H} \lesssim 10^{-7}. \tag{19}$$

We have shown that $v_\Delta = \mathcal{O}(1\,\text{GeV})$ is required to explain the CDF measurement of W boson mass, which naturally leads to $(Y_D)^{1k} \lesssim 10^{-4}$ for $\mu_H \sim \tilde{m} = \mathcal{O}(1\,\text{TeV})$. Thus, using the type III seesaw formula, one can roughly estimate the size of $(Y_D)^{1k}$ as

$$(Y_D)^{1k} \sim \frac{\sqrt{m_\nu m_\Delta}}{v_u}, \tag{20}$$

where $m_\nu \sim 10^{-10}$ GeV is the neutrino mass scale and $v_u \sim v = 246$ GeV is the electroweak VEV. We find that the bound on the Yukawa coupling $(Y_D)^{1k} \lesssim 10^{-4}$ is satisfied for

$$m_\Delta \lesssim 6 \times 10^6 \text{ GeV}. \tag{21}$$

It is worth pointing out that the left-handed sneutrinos ($\widetilde{\nu}_i$) in the lepton doublet superfields can develop induced VEVs, $\widetilde{\nu}_i \sim \varepsilon_i v_d$, through their mixing with H_d. This generates neutrino masses, shown in Fig. 1 (b), and the mass given by

$$m_\nu \sim \frac{g^2 \langle \widetilde{\nu}_i \rangle^2}{m_{\widetilde{W}}}, \tag{22}$$

where $m_{\widetilde{W}}$ is the wino mass. For $m_{\widetilde{W}} = \tilde{m} \simeq \mathcal{O}(1\,\text{TeV})$ and $\varepsilon_i \lesssim 10^{-7}$, we obtain

$$m_\nu^{\text{new}} \sim \frac{\left(10^{-7} v_d\right)^2}{m_{\widetilde{W}}} \sim 10^{-13} \text{ GeV}, \tag{23}$$

which is much smaller than the typical neutrino mass scale of 10^{-10} GeV required by the neutrino oscillation data [54]. Therefore, the R-parity violation has no significant effect on the type III seesaw mechanism.
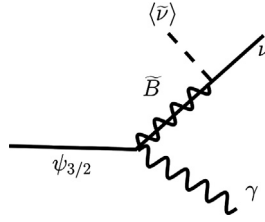
Fig. 3. Gravitino DM decay into neutrino and a photon ($\nu + \gamma$).

4. Gravitino dark matter

In MSSM, a stable neutralino is a standard WIMP candidate for DM. Since R-parity is violated in our model, the neutralino is not stable and hence it is not a viable DM candidate. We show how a long-lived gravitino is a viable DM candidate. Let us consider a gravitino with mass in the few GeV range which mainly decays to $\gamma + \nu$ [55] as shown in the Fig. 3. Its decay width is approximately given by

$$\Gamma_{\psi_{3/2} \to \gamma \nu} \sim \frac{1}{16\pi m_{3/2}} \left| \frac{(m_{3/2})^3}{\langle F \rangle} \right|^2 \left| \frac{g_Y \langle \tilde{\nu}_i \rangle}{m_{\tilde{B}}} \right|^2 , \qquad (24)$$

where $\langle F \rangle$ is the SUSY breaking order parameter, $m_{3/2} \sim \langle F \rangle / M_p$ is the gravitino mass, and $M_p = 2.44 \times 10^{18}$ is the reduced Planck mass. With $m_\nu^{\text{new}} = g_Y \langle \tilde{\nu} \rangle \sim 10^{-13}$ GeV, $m_{3/2} = \mathcal{O}(1)$ GeV and $m_{\tilde{B}} = \mathcal{O}(1)$ TeV, we obtain $\tau_{3/2} \sim 10^{30}$ seconds for the gravitino lifetime, which makes it a viable candidate for (non-thermal) cold DM. For a discussion of gravitino DM scenarios in the context of type III seesaw extended MSSM, see Ref. [56], where the gauge mediated SUSY breaking scenario is unified with the type III seesaw mechanism. Also, see Ref. [57] for the discussion of a successful inflation scenario implementation to the model of Ref. [56]. Since the inclusion of the R-parity violating term does not alter the underlying physics of these scenarios, we can realize both of these scenarios in our model. Finally, it is straightforward to embed our scenario in a GUT framework such as $SU(5)$ (see, for example, Ref. [56]).

5. Conclusion

To summarize, we have presented an extension of MSSM which incorporates type III neutrino seesaw and also includes an appropriate R-parity violating term. This leads to a tree-level correction that increases the W boson mass, such that a 2-σ agreement both with the recent determination of m_W by CDF as well as the LEP measurement of the SM ρ parameter is achieved (see Fig. 2). Based on some recent papers, we expect to be able to reach the central CDF value for m_W by taking into account radiative corrections involving some of the supersymmetric particles. Since R-parity is not conserved, we identified a few GeV long-lived gravitino as a plausible candidate for (non-thermal) cold dark matter. Finally, our model can be readily embedded in a grand unified framework such as $SU(5)$.

CRediT authorship contribution statement

Equal Contribution by all the authors and the author names are arranged alphabetically.

Declaration of competing interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

Data availability

No data was used for the research described in the article.

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