

Study of $e^+e^- \rightarrow \Omega^-\bar{\Omega}^+$ at center-of-mass energies from 3.49 to 3.67 GeV

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Using data samples of e^+e^- collisions collected with the BESIII detector at eight center-of-mass energy points between 3.49 and 3.67 GeV, corresponding to an integrated luminosity of 670 pb^{-1} , we present the upper limits of Born cross sections and the effective form factor for the process $e^+e^- \rightarrow \Omega^-\bar{\Omega}^+$. A fit to the cross sections using a perturbative QCD-derived energy-dependent function shows no significant threshold effect. The upper limit on the measured effective form factor is consistent with a theoretical prediction within the uncertainty of 1σ . These results provide new experimental information on the production mechanism of Ω .

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I. INTRODUCTION

The Ω^- resonance, composed of three valence strange quarks, is one of the most famous hyperons. Its discovery was guided directly by the eightfold way [1] more than half a century ago. Until now, its physical properties and inner structure are still not well understood. One approach to parametrize its inner structure is via electromagnetic form factors which can be accessed experimentally with electron-positron annihilation into a virtual photon. For a spin- S baryon, there are $2S + 1$ form factors to describe the γ^* -baryon-antibaryon vertex in the electron-positron annihilation under the assumption of one photon exchange. For baryons with $S = \frac{1}{2}$, the two form factors are the magnetic form factor $|G_M(q^2)|$ and the electric form factor $|G_E(q^2)|$. For baryons with $S = \frac{3}{2}$, like the Ω^- hyperon, there are four form factors to describe the $\gamma^*\Omega^-\bar{\Omega}^+$ vertex, corresponding to electric charge ($|G_{E0}|$), magnetic dipole ($|G_{M1}|$), electric quadrupole ($|G_{E2}|$), and magnetic octupole ($|G_{M3}|$), respectively [2]. The individual form factors can be determined through the analysis of the angular distribution of the hyperon in the electron-positron c.m. frame. Due to limited statistics, most experiments have measured only a combination of the electromagnetic form factors, the so-called effective form factor $|G_{\text{eff}}|$, which can be extracted from the total cross section of the pair production in electron-positron annihilation [3–8]. A recent calculation of the effective form factor has been performed for the Ω^- hyperon using the covariant spectator quark model [9], in which the data from the CLEO-c experiment [10] in the timelike region were used to fix the free parameters.

In addition, many experimental studies determining the cross section of baryon-antibaryon pairs observed an unusual behavior near the threshold. The cross sections for the production of neutron [11–13], proton [14–17], Λ [18], and Λ_c pairs [19] are approximately constant. The

only data for the process $e^+e^- \rightarrow \Omega^-\bar{\Omega}^+$ were gathered by the CLEO-c experiment at $q^2 \geq 14.2 \text{ GeV}^2$ [10], which is far from the threshold of $\Omega^-\bar{\Omega}^+$ at $q_{\text{thr}}^2 = 11.19 \text{ GeV}^2$.

The nonvanishing cross section near threshold and the wide-range plateau have attracted great interest and driven a lot of theoretical studies. Some of the explanations put forward include a $B\bar{B}$ bound state or unobserved meson resonances [20], Coulomb final-state interactions or quark electromagnetic interaction, and scenarios that take into account the asymmetry between attractive and repulsive Coulomb factors [21]. However, no obvious threshold effect was observed in the reactions $e^+e^- \rightarrow \Sigma\bar{\Sigma}$, $\Xi^-\bar{\Xi}^+$, and $\Xi^0\bar{\Xi}^0$ [22–24], indicating that the baryon antibaryon pair production in e^+e^- annihilation near the threshold region is not fully understood.

In this article, we report a measurement of the Born cross section and the effective form factor for the process $e^+e^- \rightarrow \Omega^-\bar{\Omega}^+$ based on data samples collected at eight c.m. energies $\sqrt{s} = 3.4900, 3.5080, 3.5097, 3.5104, 3.5146, 3.5815, 3.6500, \text{ and } 3.6702 \text{ GeV}$ with the BESIII detector [25] at the BEPCII [26].

II. BESIII DETECTOR AND DATA SAMPLES

The BESIII detector [25] records symmetric e^+e^- collisions provided by the BEPCII storage ring [26] in the center-of-mass energy range from 2.0 to 4.95 GeV [27], with a peak luminosity of $1 \times 10^{33} \text{ cm}^{-2} \text{ s}^{-1}$ achieved at $\sqrt{s} = 3.77 \text{ GeV}$. The cylindrical core of the BESIII detector covers 93% of the full solid angle and consists of a helium-based multilayer drift chamber (MDC), a plastic scintillator time-of-flight (TOF) system, and a CsI(Tl) electromagnetic calorimeter (EMC), which are all enclosed in a superconducting solenoidal magnet providing a 1.0 T magnetic field. The solenoid is supported by an octagonal flux-return yoke with resistive plate counter muon identification modules interleaved with steel [28]. The charged-particle momentum resolution at 1 GeV/c is 0.5%, and the specific energy loss (dE/dx) resolution is 6% for electrons from Bhabha scattering. The EMC measures photon energies with a resolution of 2.5% (5%) at 1 GeV in the barrel (end cap) region. The time resolution in the TOF

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barrel region is 68 ps, while that in the end cap region is 110 ps. The end cap TOF system was upgraded in 2015 using multigap resistive plate chamber technology, providing a time resolution of 60 ps [29].

To determine the selection efficiency, 200,000 $e^+e^- \rightarrow \Omega^-\bar{\Omega}^+$ signal MC events are generated evenly distributed in the phase space (PHSP) for each of the eight energy points using the generator CONEXC [30], which takes into account the beam-energy spread and corrections from initial-states radiation (ISR). The Ω^- ($\bar{\Omega}^+$) hyperon is modeled with EvtGen [31] to decay inclusively according to the branching fractions taken from the Particle Data Group (PDG) [32]. The response of the BESIII detector is modeled using a framework based on GEANT 4 [33]. Large simulated samples of generic $e^+e^- \rightarrow$ hadrons events (inclusive MC) are used to study possible background reactions with a generic event type analysis tool, TopoAna [34].

III. EVENT SELECTION

As the selection of $e^+e^- \rightarrow \Omega^-\bar{\Omega}^+$ events with a full reconstruction of all six final-state particles suffers from a low reconstruction efficiency, a single hyperon tag technique is applied and the events are reconstructed in one of three categories: (1) Only the Ω^- is reconstructed through the decay chain of $\Omega^- \rightarrow \Lambda K^-$ and $\Lambda \rightarrow p\pi^-$. The non-reconstructed $\bar{\Omega}^+$ is identified in the recoil against Ω^- . (2) Only the $\bar{\Omega}^+$ is reconstructed and the Ω^- is missed. (3) Both Ω^- and $\bar{\Omega}^+$ are fully reconstructed. The three categories are mutually exclusive and there is no double counting for the reconstructed events.

Charged tracks are required to be reconstructed within the angular coverage of the MDC: $|\cos\theta| < 0.93$, where θ is the polar angle with respect to the direction of the positron beam. The dE/dx information obtained from MDC combined with the flight time in the TOF is used to calculate the probability for the track being a pion, kaon, or proton, respectively. Each track is assigned to the particle type with the highest probability. Events with at least one negatively charged pion, one negatively charged kaon, and one proton are kept for further analysis.

The intermediate Λ candidate is reconstructed by a secondary vertex fit [35] that is applied to all $p\pi^-$ combinations. The $p\pi^-$ invariant mass ($M_{p\pi^-}$) of the selected candidate is required to be within 4 MeV/ c^2 of the nominal Λ mass from the PDG [32], determined by 3σ of the Λ peak from MC studies. To further suppress background from non- Λ events, the Λ decay length is required to be greater than zero. Here, the Λ decay length is defined by the position vector from the primary to the secondary vertex, taking into account the crossing angle of the position vector with the momentum of Λ . After the above mentioned selection criteria, it is possible that multiple Λ candidates are reconstructed in one event; the

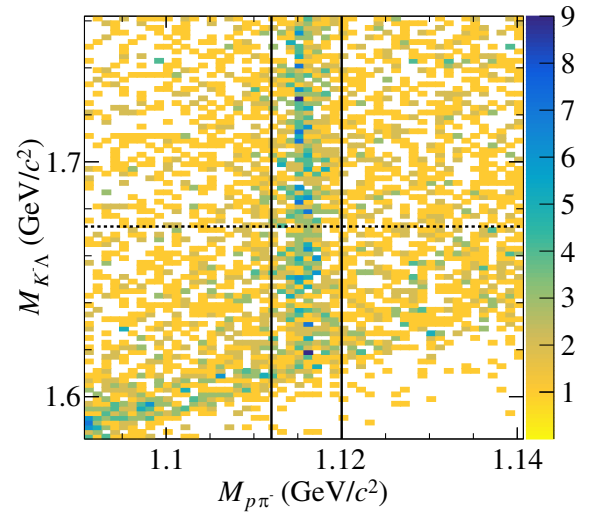


FIG. 1. Two-dimensional distribution of $M_{K^- \Lambda}$ versus $M_{p\pi^-}$ from the sum of the eight energy points. The black solid lines indicate the Λ window requirement, and the black dotted line indicates the Ω mass.

best candidate selection is performed simultaneously to the Ω reconstruction as described in the following.

The Ω^- candidate is reconstructed by combining a Λ candidate with a K^- . Similar to the Λ reconstruction, a secondary vertex fit is applied. In cases with more than one Λ or Ω^- candidate, the variable $\delta = (M_{p\pi^-} - m_\Lambda)^2 + (M_{K^- \Lambda} - m_{\Omega^-})^2$ is minimized to choose the best combination, where $M_{K^- \Lambda/p\pi^-}$ is the invariant mass of the $K^- \Lambda/p\pi^-$ pair, and m_{Λ/Ω^-} is the mass of the Λ/Ω^- hyperon taken from the PDG [32]. To suppress background events, the decay length of Ω^- is also required to be larger than zero. Figure 1 shows the distribution of $M_{K^- \Lambda}$ versus $M_{p\pi^-}$ for the combined eight energy points. A clear accumulation is observed around the mass of the Λ .

To select $\bar{\Omega}^+$ hyperon candidates, we use the distribution of the mass recoiling against the selected $K^- \Lambda$ system,

$$RM_{K^- \Lambda} = \sqrt{(\sqrt{s} - E_{K^- \Lambda})^2 - |\vec{p}_{K^- \Lambda}|^2}, \quad (1)$$

where $E_{K^- \Lambda}$ and $\vec{p}_{K^- \Lambda}$ are the energy and momentum of the selected $K^- \Lambda$ candidate in the c.m. system.

To improve the resolution, a correction is performed with $M_{K^- \Lambda}^{\text{correction}} = M_{K^- \Lambda} - M_{p\pi^-} + m_\Lambda$. Similarly, a correction is performed to the recoil side with $RM_{K^- \Lambda}^{\text{correction}} = RM_{K^- \Lambda} + M_{K^- \Lambda} - m_{\Omega^-}$. Figure 2 shows the distributions of the corrected recoil mass versus the corrected mass of Ω for each energy point, the central red box indicates the signal zone (zone S), while B_1, B_2, B_3, B_4 marked by dashed blue boxes are for the considered sideband regions.

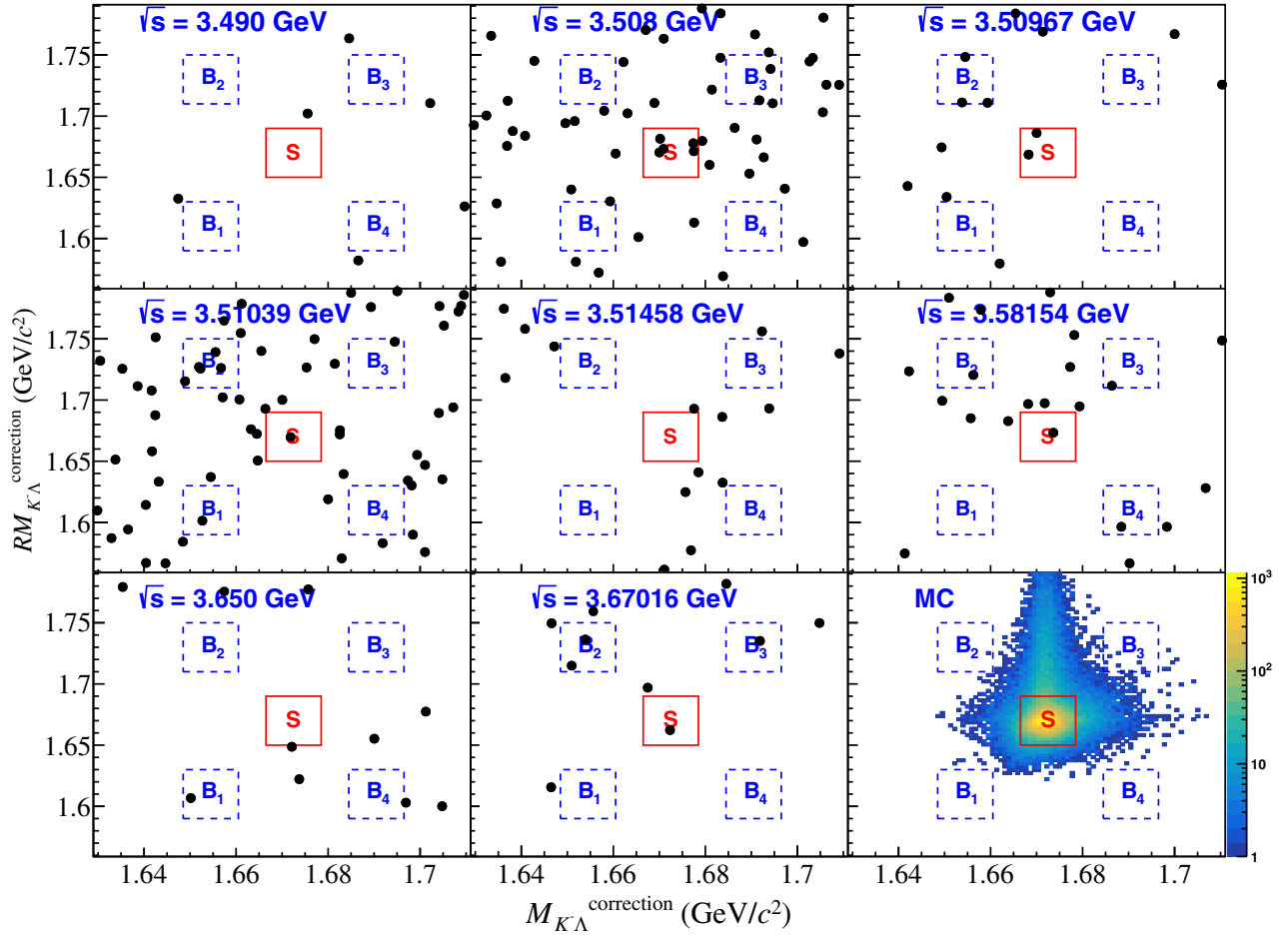


FIG. 2. Distribution of $RM_{K\Lambda}^{\text{correction}}$ versus $M_{K\Lambda}^{\text{correction}}$ for eight energy points from data and the signal MC sample at 3.490 GeV (lower right plot). The red box denotes the Ω^- signal region, while the dashed blue boxes denote the sideband regions.

IV. EXTRACTION OF SIGNAL YIELDS

The signal yield for each energy point is extracted by counting the number of events in the signal and the sideband regions as shown in Fig. 2. The signal and background regions are defined as follows:

- (i) S : $M_{\Lambda K} \in [1.6665, 1.6785]$, $RM_{\Lambda K} \in [1.65, 1.69]$.
- (ii) B_1 : $M_{\Lambda K} \in [1.6485, 1.6605]$, $RM_{\Lambda K} \in [1.59, 1.63]$.
- (iii) B_2 : $M_{\Lambda K} \in [1.6485, 1.6605]$, $RM_{\Lambda K} \in [1.71, 1.75]$.
- (iv) B_3 : $M_{\Lambda K} \in [1.6845, 1.6965]$, $RM_{\Lambda K} \in [1.71, 1.75]$.
- (v) B_4 : $M_{\Lambda K} \in [1.6845, 1.6965]$, $RM_{\Lambda K} \in [1.59, 1.63]$.

Here, the window size for the tag and the recoil side are chosen as 3σ of the mass spectrum from MC studies. The number of observed events in the signal region N_{obs} is determined. The number of background events in the sideband regions N_{bkg} is obtained by $N_{\text{bkg}} = \frac{1}{4} \sum_{i=1}^4 N_{B_i}$, where i runs over the four regions shown in Fig. 2. The number of signal events is obtained by $N_S = N_{\text{obs}} - N_{\text{bkg}}$. In Table I, the uncertainty of N_S is calculated using the Feldman-Cousins method [36]. Note that negative N_S have been set to be 0 to avoid an unphysical number of signal events. Due to the low statistics, upper limits for these

measurements are provided. The N^{UL} is the upper limit at the 90% confidence level, which is calculated with a profile likelihood method [37,38]. To investigate the potential background reactions, a study of the inclusive MC at 3.650 and 3.686 GeV is performed. Main background events are from non- Ω processes, such as $e^+e^- \rightarrow \Lambda\bar{\Lambda}\phi$ with $\phi \rightarrow K^+K^-$. These background events can be estimated using the sideband strategy.

V. DETERMINATION OF BORN CROSS SECTION

The Born cross section is calculated by

$$\sigma^B(s) = \frac{N_S}{\mathcal{L} \cdot (1 + \delta) \cdot \frac{1}{|1 - \Pi|^2} \cdot \varepsilon}, \quad (2)$$

where \mathcal{L} is the integrated luminosity, $(1 + \delta)$ is the ISR correction factor, $\frac{1}{|1 - \Pi|^2}$ is the vacuum polarization (VP) correction factor, and ε is the selection efficiency. The VP correction factor is obtained using the calculation described in Ref. [41]. The ISR correction and efficiency ε are

TABLE I. The numerical results for $e^+e^- \rightarrow \Omega^-\bar{\Omega}^+$. \mathcal{L} is the integrated luminosity [39,40], $\frac{1}{|1-\prod|^2}$ is the VP correction factor, $1 + \delta$ is the ISR correction factor, ϵ is the selection efficiency, N_{obs} denotes the number of observed events in the signal region, N_{bkg} denotes the number of background events estimated with the sideband regions, N_S (N_S^{UL}) is the number (upper limit) of signal events, σ^B represents the Born cross section, and $G_{\text{eff}}(s)$ is the effective form factor. For the cross section and form factor, the first uncertainty is statistical and the second is systematic.

\sqrt{s} (GeV)	\mathcal{L} (pb $^{-1}$)	$\frac{1+\delta}{ 1-\prod ^2}$	ϵ	N_{obs}	N_{bkg}	N_S (N_S^{UL})	σ^B (fb)	$ G_{\text{eff}}(s) (\times 10^{-3})$
3.4900	12.11	0.88	0.071	0	0	$0.0_{-0.0}^{+1.3}$ (< 2.0)	$0_{-0}^{+1780} \pm 0$ (< 2738)	$0_{-0}^{+24} \pm 0$ (< 30)
3.5080	181.79	0.89	0.075	5	0.5	$4.5_{-2.2}^{+2.8}$ (< 9.6)	$371_{-185}^{+232} \pm 16$ (< 797)	$11_{-3}^{+3} \pm 1$ (< 16)
3.5097	39.29	0.89	0.078	2	0.75	$1.3_{-0.9}^{+2.3}$ (< 5.3)	$458_{-339}^{+825} \pm 20$ (< 1947)	$12_{-6}^{+8} \pm 1$ (< 25)
3.5104	183.64	0.89	0.077	1	1.5	$0.0_{-0.0}^{+1.3}$ (< 2.6)	$0_{-0}^{+105} \pm 0$ (< 231)	$0_{-0}^{+6} \pm 0$ (< 9)
3.5146	40.92	0.89	0.080	0	0	$0.0_{-0.0}^{+1.3}$ (< 2.0)	$0_{-0}^{+443} \pm 0$ (< 682)	$0_{-0}^{+12} \pm 0$ (< 14)
3.5815	85.28	0.92	0.100	1	0.75	$0.3_{-0.3}^{+1.8}$ (< 3.6)	$32_{-32}^{+225} \pm 1$ (< 469)	$3_{-3}^{+6} \pm 1$ (< 11)
3.6500	44.49	0.92	0.120	0	0.25	$0.0_{-0.0}^{+1.0}$ (< 2.0)	$0_{-0}^{+214} \pm 0$ (< 408)	$0_{-0}^{+7} \pm 0$ (< 10)
3.6702	83.61	0.90	0.120	1	0.75	$0.3_{-0.3}^{+1.8}$ (< 3.6)	$29_{-29}^{+200} \pm 1$ (< 417)	$3_{-3}^{+5} \pm 1$ (< 10)

determined using an iterative approach where a flat cross section line shape is adopted as an initial input and is iterated to obtain a stable result. In this analysis, the last two iterations have a difference of 0.03% on the cross section. The measured Born cross sections of the eight energy points are listed in Table I.

VI. DETERMINATION OF EFFECTIVE FORM FACTOR

For the spin- $\frac{3}{2}$ Ω^- hyperon, the effective form factor $|G_{\text{eff}}(s)|$ can be defined by a combination of the four form factors: $|G_{E0}|$, $|G_{M1}|$, $|G_{E2}|$, and $|G_{M3}|$ [2,9].

$$|G_{\text{eff}}(s)| = \sqrt{\frac{2 \times \frac{s}{4m^2} |G_M^*(s)|^2 + |G_E^*(s)|^2}{2 \times \frac{s}{4m^2} + 1}}. \quad (3)$$

Here, m is the Ω^- mass, and $|G_E^*(s)|$ and $|G_M^*(s)|$ are defined as follows [9]:

$$\begin{aligned} |G_E^*(s)|^2 &= 2|G_{E0}|^2 + \frac{8}{9} \left(\frac{s}{4m^2} \right)^2 |G_{E2}|^2, \\ |G_M^*(s)|^2 &= \frac{10}{9} |G_{M1}|^2 + \frac{32}{5} \left(\frac{s}{4m^2} \right)^2 |G_{M3}|^2. \end{aligned} \quad (4)$$

The total cross section of $e^+e^- \rightarrow \Omega^-\bar{\Omega}^+$ is related to the form factors as

$$\sigma^B(s) = \frac{4\pi\alpha^2 C\beta}{3s} \left[|G_M^*(s)|^2 + \frac{2m^2}{s} |G_E^*(s)|^2 \right], \quad (5)$$

where α is the fine structure constant, $\beta = \sqrt{1 - \frac{4m^2}{s}}$ is the velocity, and $C = y/(1 - e^{-y})$ is the Coulomb factor parametrizing the electromagnetic interaction between the outgoing baryon and antibaryon with $y = \pi \cdot \alpha \cdot \sqrt{(1 - \beta^2)}/\beta$.

Using Eqs. (3) and (5), the effective form factor can be determined from the total cross section by

$$|G_{\text{eff}}(s)| = \sqrt{\frac{\sigma^B(s)}{(1 + \frac{2m^2}{s}) \cdot (\frac{4\pi\alpha^2 C\beta}{3s})}}. \quad (6)$$

For the eight energy points in this analysis, the corresponding effective form factors are listed in Table I. The upper limits on the cross sections and form factors are determined using the profile likelihood method incorporating the systematic uncertainties, where the systematic uncertainties are also included in the profile likelihood method.

VII. FIT TO BORN CROSS SECTION

A least- χ^2 fit to the Born cross section of $e^+e^- \rightarrow \Omega^-\bar{\Omega}^+$ is performed with a perturbative QCD (pQCD) driven energy power law function,

$$\sigma^B(\sqrt{s}) = \frac{c_0 \cdot \beta \cdot C}{(\sqrt{s} - c_1)^{10}}, \quad (7)$$

where c_0 and c_1 are free parameters. This model has been applied in studies of $e^+e^- \rightarrow \Lambda\bar{\Lambda}$ [18], $\Xi^0\bar{\Xi}^0$ [23], and $\Xi^-\bar{\Xi}^+$ [24] production. The fit returns $c_0 = (74.8_{-73.0}^{+3103.1})$ pb \cdot GeV $^{-10}$ and $c_1 = (2.65 \pm 0.17)$ GeV, where both statistical and systematic contributions of the measured cross section have been included in the fit without considering the correlation between different energy points. Figure 3 shows the fit result with quality $\chi^2/n.d.o.f. = 4.2/6.0$. The data can be well described by the pQCD driven energy power function, which indicates no obvious threshold enhancement.

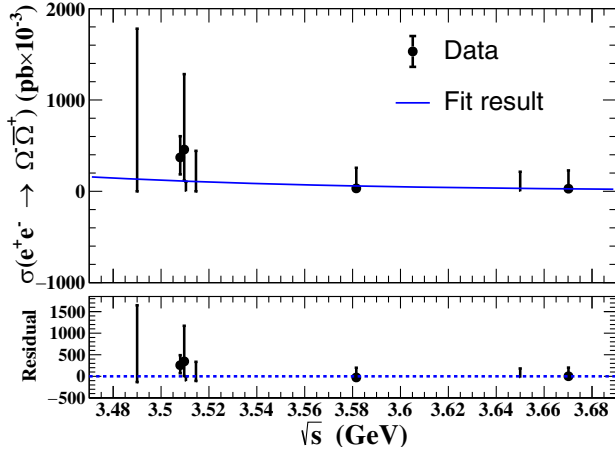


FIG. 3. Fit to the measured Born cross section with a pQCD driven energy power function. The dots with the error bars are the measured Born cross section at c.m. energies between 3.49 and 3.67 GeV. The blue line denotes the fit results. The bottom plot shows the residual distribution.

VIII. SYSTEMATIC UNCERTAINTY

Several sources of systematic uncertainties are considered on the Born cross section measurement. They include the Ω reconstruction efficiency, the mass windows and decay length requirements of Λ and Ω , sideband selection, angular distribution, the MC generator, and the fit models for the cross section. Limited knowledge of the decay branching fractions of intermediate states and the luminosity measurement provide additional contributions. All the systematic uncertainties are discussed in detail below.

- (1) The systematic uncertainty due to the Ω^- reconstruction efficiency, including the tracking and particle identification efficiency for the charged tracks and the efficiency for the Λ reconstruction, is estimated by the control sample $\psi(3686) \rightarrow \Omega^- \bar{\Omega}^+$ with the same method as described in Refs. [42–48].
- (2) The uncertainty associated with the mass windows of Λ and Ω is estimated by changing the nominal window size from 3σ to 4σ . The difference in results is taken as the uncertainty. The uncertainty from the decay lengths cut of Λ or Ω^- is estimated by changing the nominal requirement of $L > 0$ to an alternative requirement of $L/\sigma_L > 2$. Here σ_L is the resolution of the decay length.
- (3) The nominal sideband regions are defined as the windows of $[6\sigma, 12\sigma]$ from the central value. By varying the position of the sideband window to $[4.5\sigma, 10.5\sigma]$ or $[7.5\sigma, 13.5\sigma]$, the maximum difference on the results is taken as the systematic uncertainty.

- (4) In this analysis, the selection efficiency for $e^+e^- \rightarrow \Omega^- \bar{\Omega}^+$ is determined based on a PHSP model, which may differ from the real angular distribution. Due to the limited statistics, it is difficult to perform a study of the angular distribution in detail. Alternatively, we utilize the angular distribution from theoretical prediction near threshold [49] to reproduce a MC sample and take the efficiency difference between the signal MC samples and the alternative MC as the systematic uncertainty due to the Ω^- angular distribution.
- (5) The precision of the ISR calculation in the MC generator is better than 1%. Here we conservatively assign 1% as the uncertainty of the radiative correction. Using CONEXC, the input line shape is iteratively modified until the final cross section becomes stable. The last two iterations have a difference of 0.03% on the cross section, which is negligible compared to the uncertainty of the MC generator.
- (6) The fit to the line shape of the Born cross section has influence on the ISR factor and the selection efficiency. The input line shape is changed by $\pm 1\sigma$, where σ is taken from the fit result. The resulting change in ISR factor and selection efficiency is taken as the systematic uncertainty due to the fit on the line shape.
- (7) The uncertainties associated with the branching fractions of the intermediate states Ω^- and Λ are taken from the PDG [32].
- (8) The luminosity at each energy point is measured using Bhabha scattering events with an uncertainty about 0.6% [39,40], which is taken as the systematic uncertainty.

Table II summarizes the various systematic uncertainties on the cross section measurements. Assuming all sources to be independent, the total systematic uncertainty is obtained by summing over the individual contributions in quadrature.

TABLE II. Systematic uncertainty on the measurement of the Born cross section (%). The uncertainties are the same for the eight energy points.

Source	Value (%)
Ω reconstruction	3.7
Λ and Ω mass windows	0.4
Λ and Ω decay lengths	1.5
Sideband	0.3
Angular distribution	0.8
MC generator	1.0
Fit on the line shape	0.2
Intermediate states	0.9
Luminosity	0.6
Total	4.3

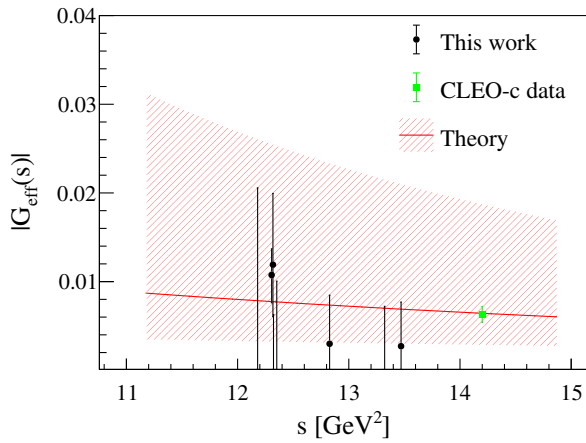


FIG. 4. Comparison of the effective form factor between this work, the CLEO-c data, and the theoretical prediction [9]. The black points with error bars are the results from this work. The green square is from a paper using the CLEO-c data [10] where $G_E = 0$ was assumed. The red band indicates the theoretical prediction with the red line in the middle showing the predicted central value [9], using CLEO-c data to fix the free parameters of the model.

IX. SUMMARY

In summary, using e^+e^- collision data corresponding to a total luminosity of 670 pb^{-1} collected with the BESIII detector at BEPCII, the upper limits on the Born cross section and effective form factor for the process $e^+e^- \rightarrow \Omega^-\bar{\Omega}^+$ are measured by means of a single hyperon tag method, at eight c.m. energies between 3.49 and 3.67 GeV. The corresponding results are listed in Table I. After a fit to the cross section of $e^+e^- \rightarrow \Omega^-\bar{\Omega}^+$ with a pQCD driven energy function, as shown in Fig. 3, no significant threshold effect is observed near the $\Omega^-\bar{\Omega}^+$ threshold. Note that the lowest energy point in this work is still about 150 MeV above the threshold.

The results of this analysis provide new experimental information to understand the production mechanism for hyperons with strangeness $S = -3$. The effective form factor determined in this paper (black data points) is consistent with the theoretical prediction (red band) using the covariant spectator quark model [9] within the uncertainty of 1σ as shown in Fig. 4. Although only upper limits on the effective form factor are obtained due to the low statistics in this measurement, they will provide a constraint

for the theoretical studies on the threshold effects of baryon pair production through e^+e^- annihilation.

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- [1] M. Gell-Mann, *Phys. Rev.* **125**, 1067 (1962).
 [2] J. G. Körner and M. Kuroda, *Phys. Rev. D* **16**, 2165 (1977).
 [3] A. Denig and G. Salme, *Prog. Part. Nucl. Phys.* **68**, 113 (2013).

- [4] S. Pacetti, R. B. Ferroli, and E. Tomasi-Gustafsson, *Phys. Rep.* **550–551**, 1 (2015).
 [5] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. Lett.* **124**, 032002 (2020).

- [6] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. D* **104**, L091104 (2021).
- [7] X. F. Wang and G. S. Huang, *Symmetry* **14**, 65 (2022).
- [8] X. F. Wang, *Rev. Mex. Fis. Suppl.* **3**, 0308074 (2022).
- [9] G. Ramalho, *Phys. Rev. D* **103**, 074018 (2021).
- [10] S. Dobbs, K. K. Seth, A. Tomaradze, T. Xiao, and G. Bonvicini, *Phys. Rev. D* **96**, 092004 (2017).
- [11] M. Ablikim *et al.* (BESIII Collaboration), *Nat. Phys.* **17**, 1200 (2021).
- [12] A. Antonelli *et al.* (FENICE Collaboration), *Nucl. Phys.* **B517**, 3 (1998).
- [13] M. N. Achasov *et al.*, *Phys. Rev. D* **90**, 112007 (2014).
- [14] J. P. Lees *et al.* (BABAR Collaboration), *Phys. Rev. D* **87**, 092005 (2013).
- [15] R. R. Akhmetshin *et al.* (CMD-3 Collaboration), *Phys. Lett. B* **759**, 634 (2016).
- [16] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. Lett.* **124**, 042001 (2020).
- [17] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Lett. B* **817**, 136328 (2021).
- [18] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. D* **97**, 032013 (2018).
- [19] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. Lett.* **120**, 132001 (2018).
- [20] O. D. Dalkarov, P. A. Khakhulin, and A. Y. Voronin, *Nucl. Phys.* **A833**, 104 (2010); B. El-Bennich, M. Lacombe, B. Loiseau, and S. Wycech, *Phys. Rev. C* **79**, 054001 (2009).
- [21] J. Haidenbauer, H.-W. Hammer, U.-G. Meißner, and A. Sibirtsev, *Phys. Lett. B* **643**, 29 (2006); R. Baldini, S. Pacetti, A. Zallo, and A. Zichichi *Eur. Phys. J. A* **39**, 315 (2009).
- [22] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Lett. B* **814**, 136110 (2021).
- [23] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Lett. B* **820**, 136557 (2021).
- [24] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. D* **103**, 012005 (2021).
- [25] M. Ablikim *et al.* (BESIII Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **614**, 345 (2010).
- [26] C. H. Yu *et al.*, in *Proceedings of IPAC2016, Busan, Korea* (JACoW, Geneva, Switzerland, 2016).
- [27] M. Ablikim *et al.* (BESIII Collaboration), *Chin. Phys. C* **44**, 040001 (2020).
- [28] K. X. Huang *et al.*, *Nucl. Sci. Tech.* **33**, 142 (2022).
- [29] X. Li *et al.*, *Radiat. Detect. Technol. Methods* **1**, 13 (2017); Y. X. Guo *et al.*, *Radiat. Detect. Technol. Methods* **1**, 15 (2017); P. Cao *et al.*, *Nucl. Instrum. Methods Phys. Res., Sect. A* **953**, 163053 (2020).
- [30] R. G. Ping, *Chin. Phys. C* **38**, 083001 (2014).
- [31] D. J. Lange, *Nucl. Instrum. Methods Phys. Res., Sect. A* **462**, 152 (2001); R. G. Ping, *Chin. Phys. C* **32**, 599 (2008).
- [32] P. A. Zyla *et al.* (Particle Data Group), *Prog. Theor. Exp. Phys.* **2020**, 083C01 (2020).
- [33] S. Agostinelli *et al.* (GEANT4 Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **506**, 250 (2003); J. Allison *et al.*, *IEEE Trans. Nucl. Sci.* **53**, 270 (2006).
- [34] X. Y. Zhou, S. Du, G. Li, and C. Shen, *Comput. Phys. Commun.* **258**, 107540 (2021).
- [35] M. Xu *et al.*, *Chin. Phys. C* **33**, 428 (2009).
- [36] G. J. Feldman and R. D. Cousins, *Phys. Rev. D* **57**, 3873 (1998).
- [37] W. Rolke, A. M. López, and J. Conrad, *Nucl. Instrum. Methods Phys. Res., Sect. A* **551**, 493 (2005).
- [38] J. Lundberg, J. Conrad, W. Rolke, and A. Lopez, *Comput. Phys. Commun.* **181**, 683 (2010).
- [39] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. Lett.* **129**, 122001 (2022).
- [40] M. Ablikim *et al.* (BESIII Collaboration), *Chin. Phys. C* **37**, 123001 (2013).
- [41] E. A. Kuraev and V. S. Fadin, *Sov. J. Nucl. Phys.* **41**, 466 (1985), <https://inspirehep.net/literature/217313>.
- [42] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. D* **87**, 032007 (2013); **87**, 059901(E) (2013).
- [43] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. D* **93**, 072003 (2016); *Phys. Lett. B* **770**, 217 (2017).
- [44] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. D* **100**, 051101 (2019).
- [45] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. D* **104**, 092012 (2021).
- [46] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. D* **105**, L011101 (2022).
- [47] M. Ablikim *et al.* (BESIII Collaboration), *J. High Energy Phys.* **06** (2022) 074.
- [48] M. Ablikim *et al.* (BESIII Collaboration), *Phys. Rev. D* **106**, L091101 (2022).
- [49] E. Perotti, G. Fäldt, A. Kupsc, S. Leupold, and J. J. Song, *Phys. Rev. D* **99**, 056008 (2019).